

Neutrino Experiments with the Fermilab Main Injector

BENE08

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Fermilab

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Outline

- The Current and Near Term Program
- Physics goals of the future program
- NO ν A : Capabilities and Status
- The US program in the global context
- LB DUSEL
- Summary and Conclusions

The Current Neutrino Program


- **8 GeV protons** from the Booster
 - Neutrinos **from** Booster Neutrino Beam (BNB)
 - **To MiniBooNE (running)**
 - **To SciBooNE (completed in August)**
- **120 GeV protons** from the Main Injector
 - Neutrinos **from** NuMI
 - **To MINOS (running)**
 - **To MINERvA (under construction)**
 - **To NOvA (beginning construction)**

The Current Neutrino Program

- **8 GeV protons** from the Booster
 - Neutrinos **from** Booster Neutrino Beam (BNB)
 - **To MiniBooNE** (running)
 - To MicroBooNE (approved, design phase)
- **120 GeV protons** from the Main Injector
 - Neutrinos **from** NuMI
 - **To MINOS** (running)
 - To ArgoNeuT (liquid argon TPC test) (installation in progress)
 - **To MINERvA** (under construction)
 - **To NOvA** (beginning construction)


The Current Neutrino Program

- 8 GeV protons from the Booster
 - Neutrinos from Booster Neutrino Beam (BNB)


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- To MiniBooNE (running)
 - To MicroBooNE (approved, design phase)

- 120 GeV protons from the Main Injector

- Neutrinos from NuMI

- 
- To MINOS (running)
 - To ArgoNeuT (liquid argon TPC test) (installation in progress)

- To MINERvA (construction)

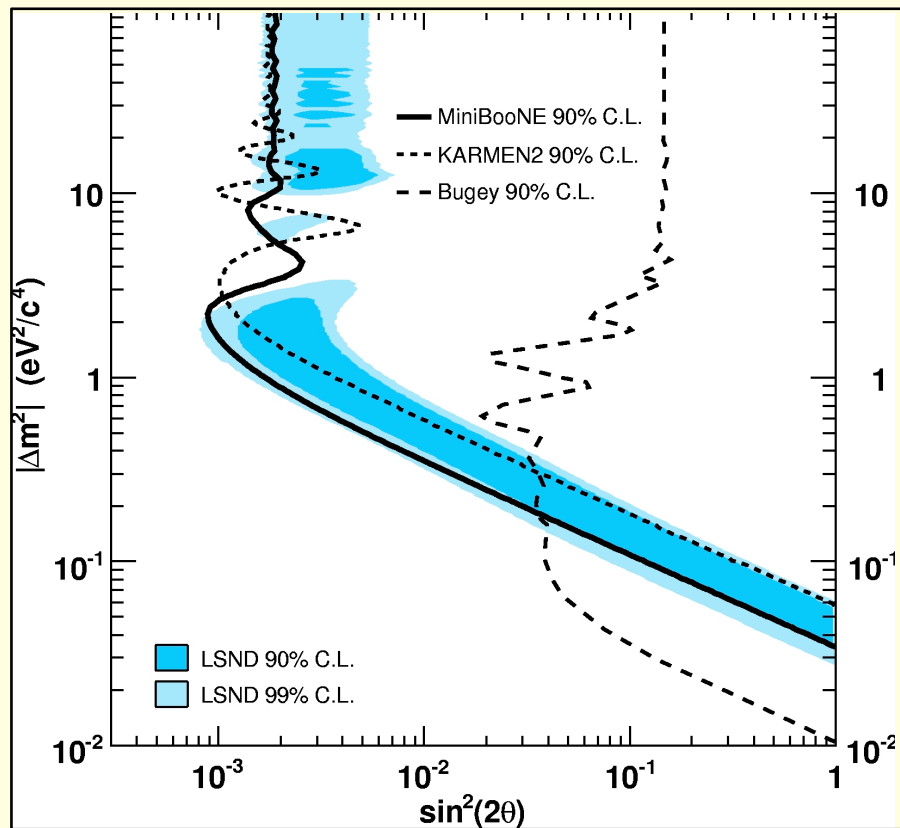
- 
- To NOvA (passed CD2 review; awaiting CD3a & funding)



Neutrino Oscillations

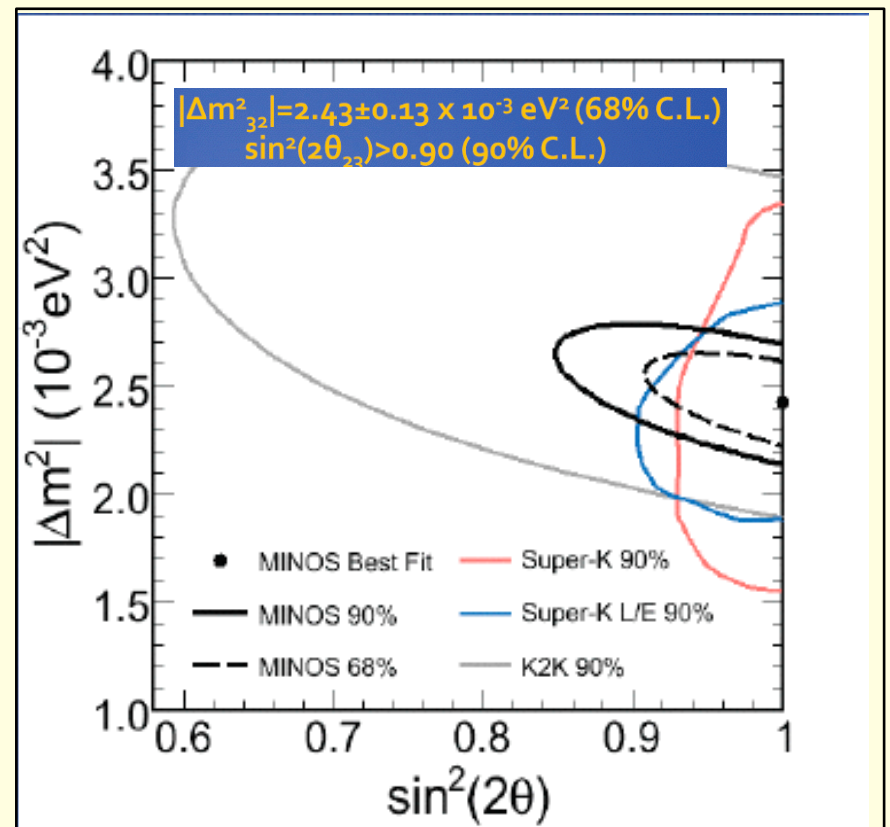
MiniBooNE $\nu_\mu \rightarrow \nu_e$ appearance

Phys. Rev. Lett. 98, 231801 (2007)



MINOS ν_μ disappearance

PRL Vol. 101, 131802 (2008)



The Quest for θ_{13}

$$\text{Flavor eigenstate} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \quad \text{Mass eigenstate}$$

At Δ_{atm} we measure the product θ_{13} and θ_{23}

$$\left\{ \begin{aligned} P_{\text{vac}}(\nu_\mu \rightarrow \nu_e) &= \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \Delta_{\text{atm}}, \\ \Delta_{\text{atm}} &\approx 1.27 \left(\frac{\Delta m_{32}^2 L}{E} \right), \end{aligned} \right.$$

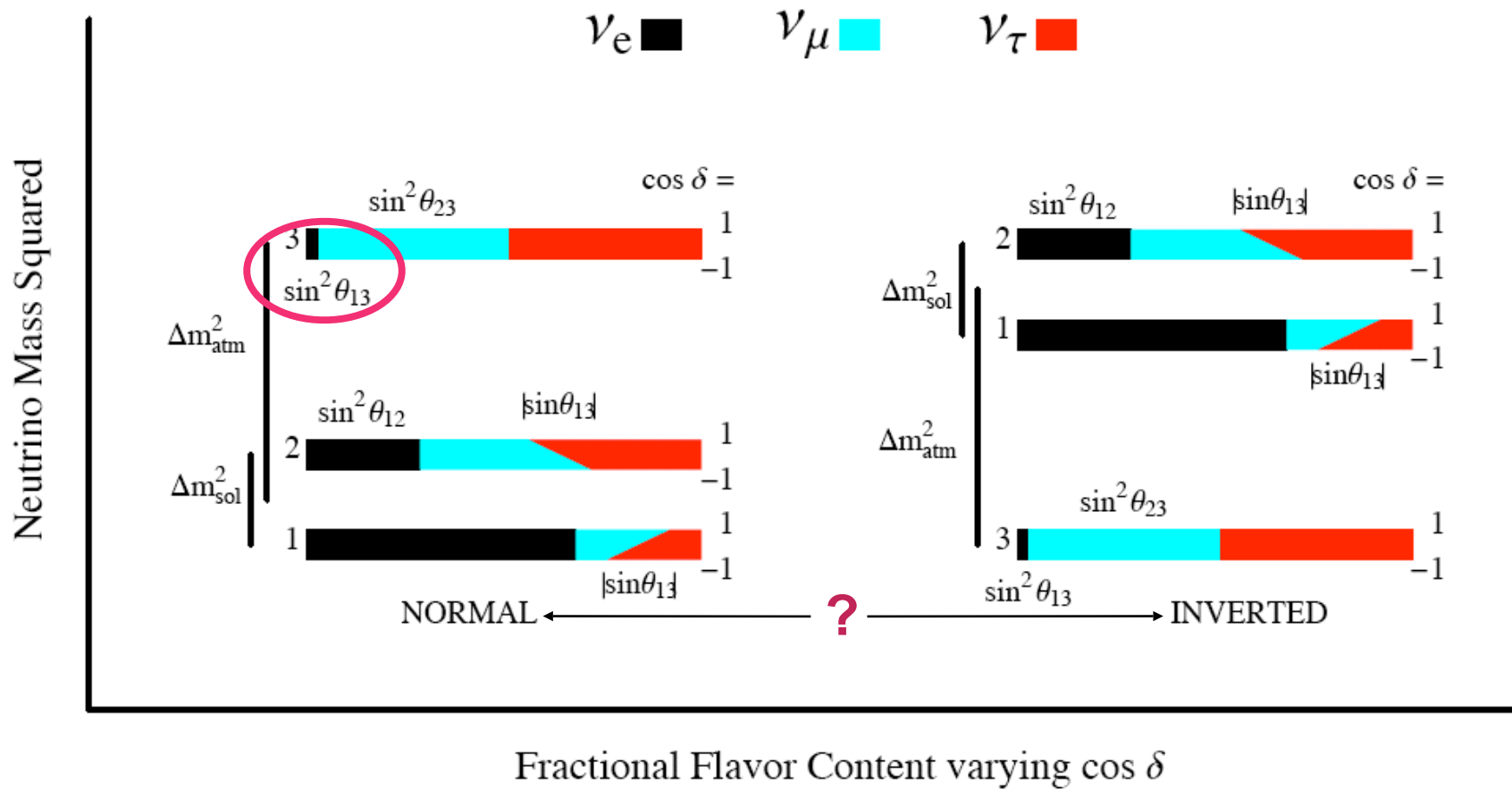
$$\left\{ \begin{aligned} P_{\text{mat}}(\nu_\mu \rightarrow \nu_e) &\approx \left(1 \pm 2 \frac{E}{E_R} \right) P_{\text{vac}}(\nu_\mu \rightarrow \nu_e) \\ E_R &= \frac{\Delta m_{32}^2}{2\sqrt{2}G_F N_e} = 12 \text{ GeV} \left(\frac{\Delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right) \left(\frac{1.4 \text{ g cm}^{-3}}{Y_e \rho} \right) \end{aligned} \right.$$

ν oscillations are enhanced, $\bar{\nu}$ are suppressed (or vice versa depending on the mass hierarchy)

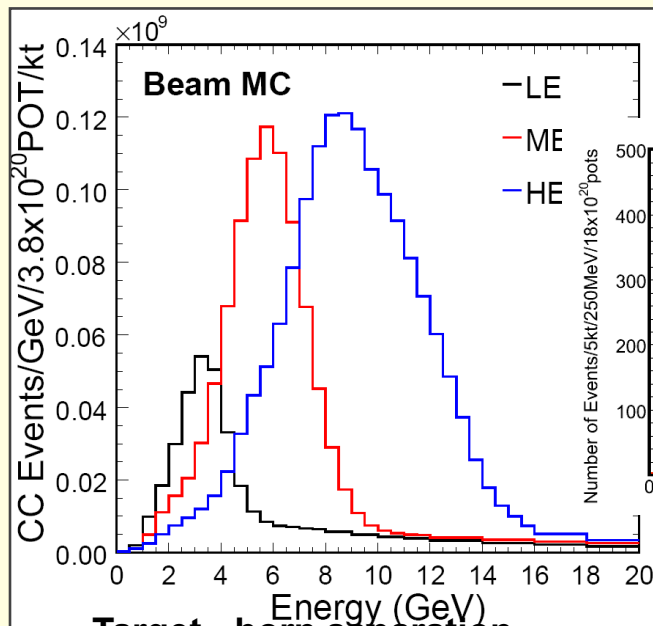
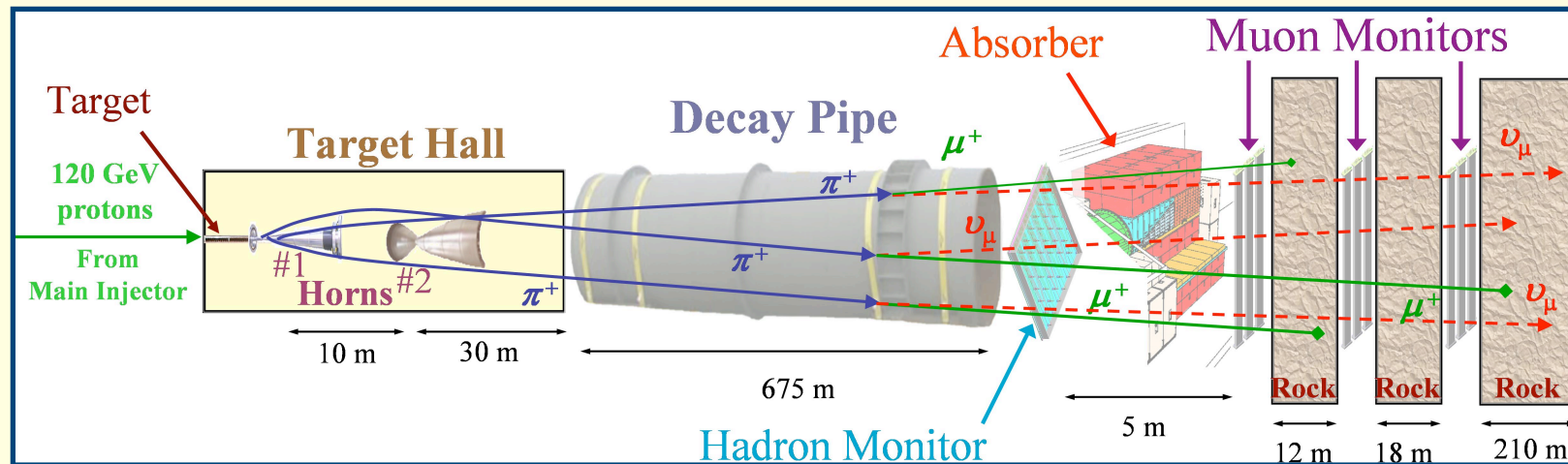
$$\left\{ \begin{aligned} \Delta P_S(\nu_\mu \rightarrow \nu_e) &\approx J_r \sin \Delta_{\text{sol}} \sin \Delta_{\text{atm}} (\cos \delta \cos \Delta_{\text{atm}} - \sin \delta \sin \Delta_{\text{atm}}), \\ J_r &= \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13}, \end{aligned} \right.$$

And the CP phase

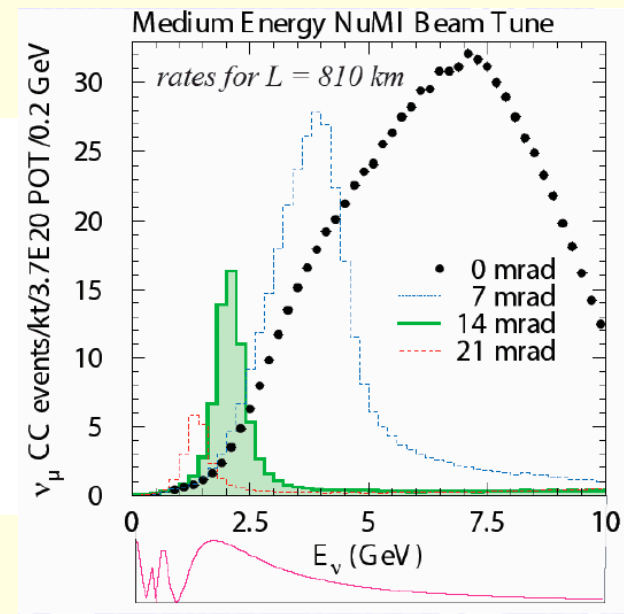
θ_{13} , mass hierarchy and δ_{CP}



The NuMI Beam



Target - horn separation sets the neutrino energy spectrum.



Off-axis detector location sees a narrow band beam

NO_vA : NuMI Off-Axis

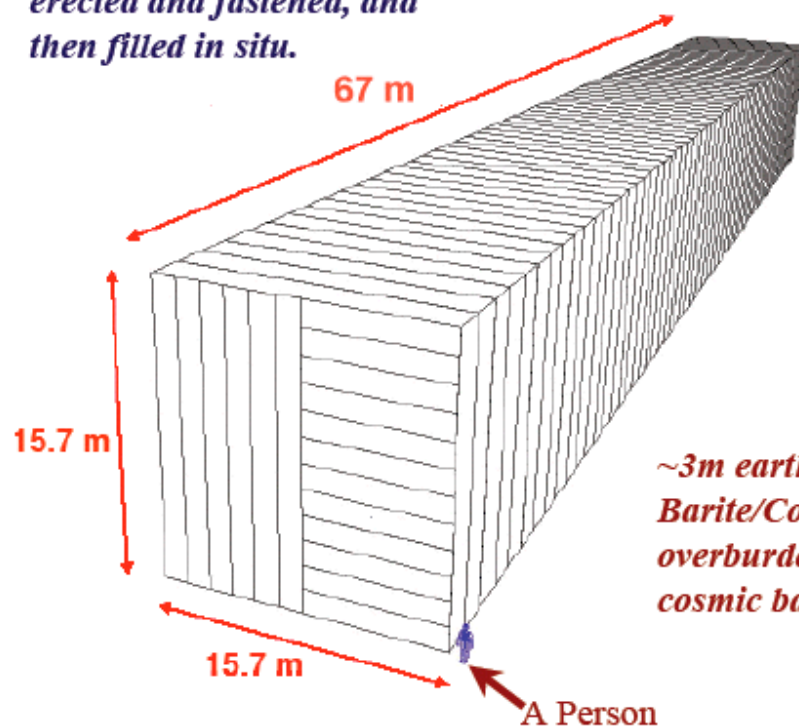




NOvA Far Detector

Construction:

Empty planes will be erected and fastened, and then filled in situ.

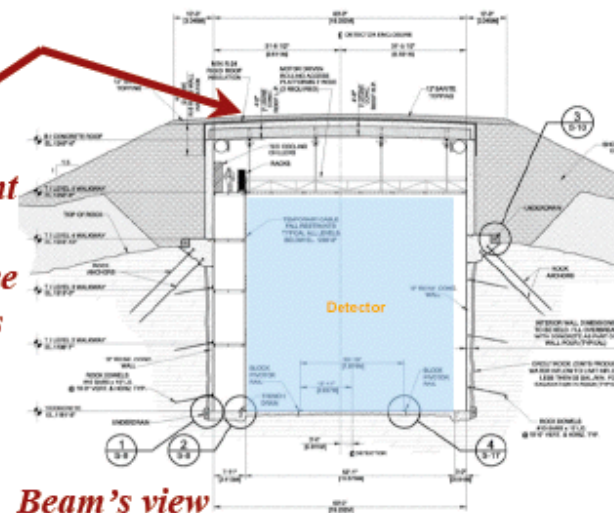


*~3m earth equivalent
Barite/Concrete
overburden to reduce
cosmic backgrounds*

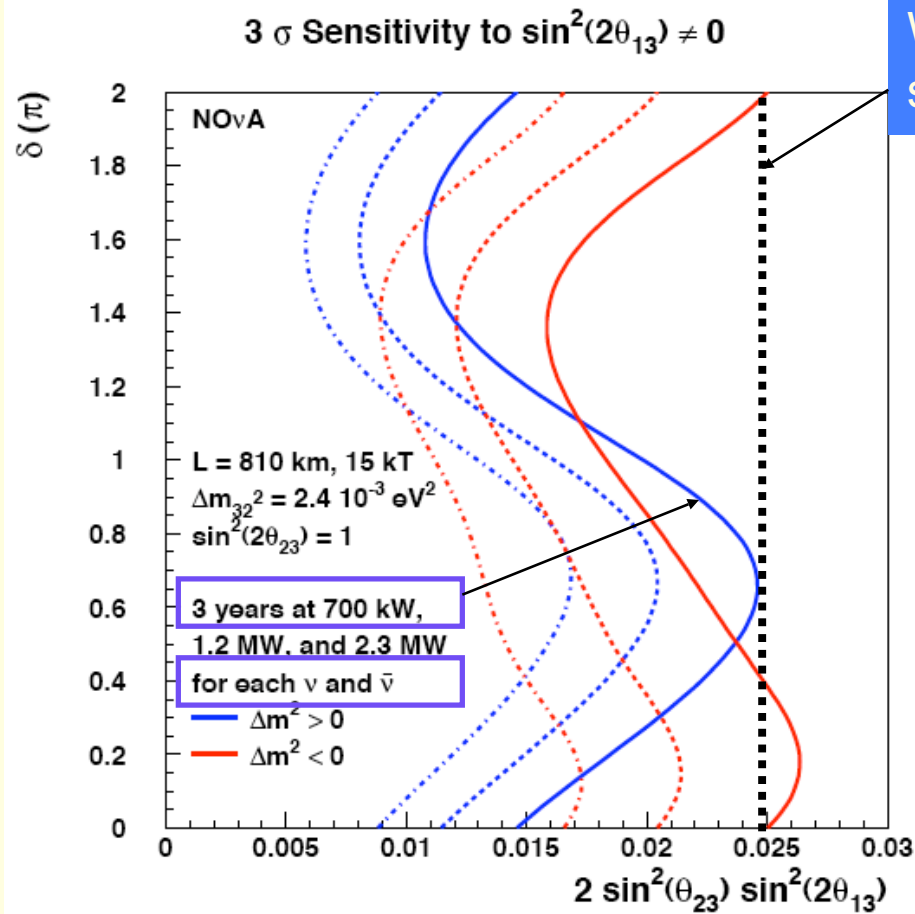
15kT Total Mass
(70% Scintillator)

1003 planes of
detector supported in
blocks of 31

385000+ cells

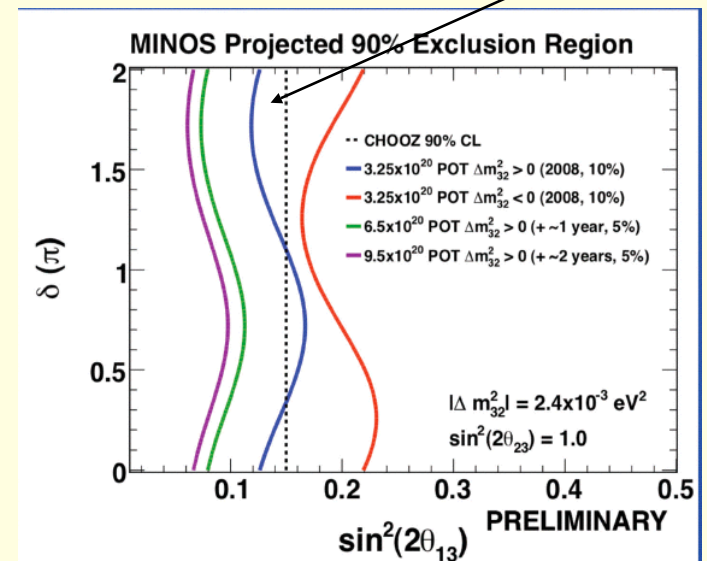


NOvA Sensitivity

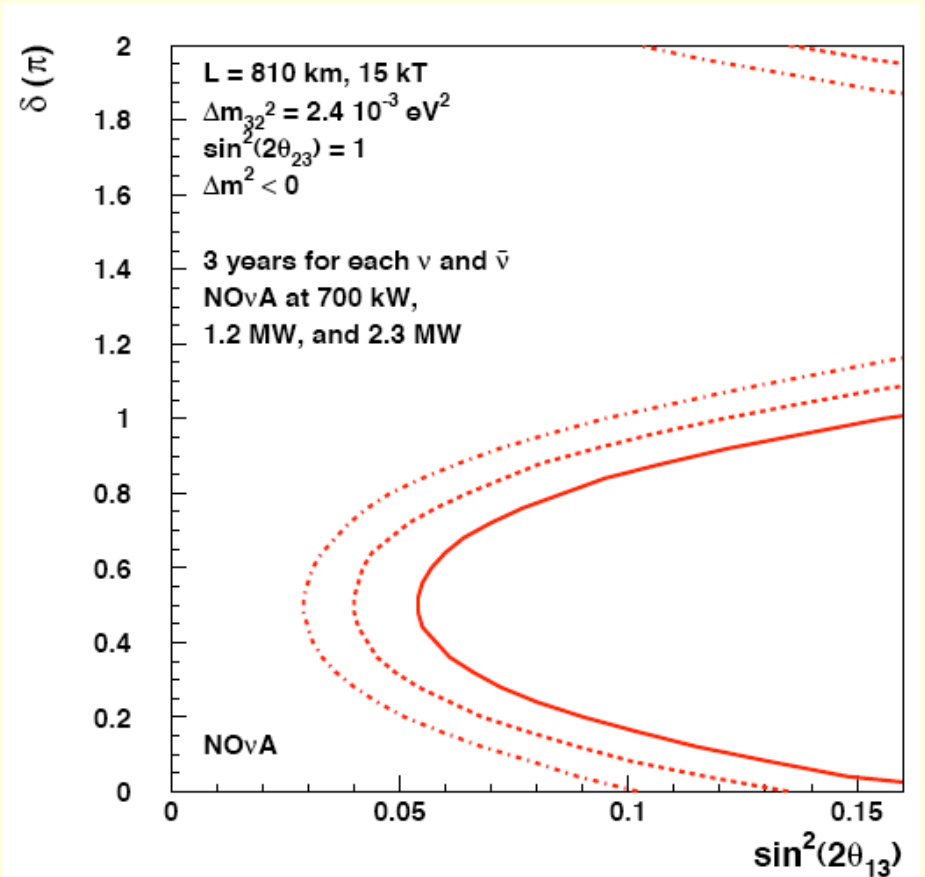
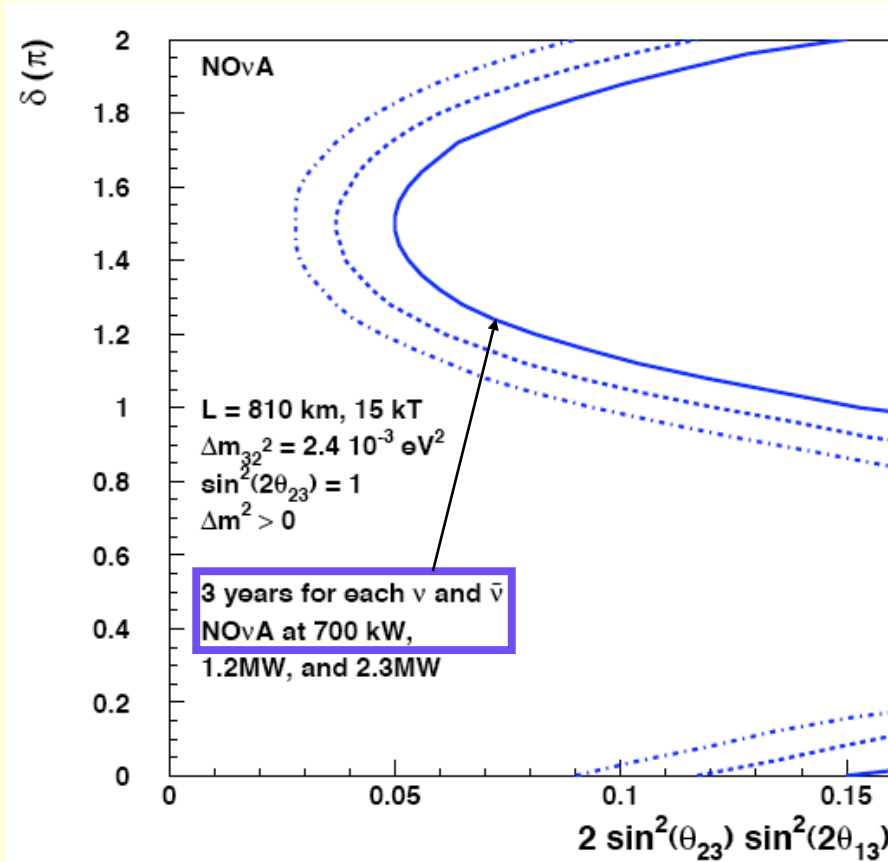


We have a $\sim 3\sigma$ discovery potential for $\sin^2 2\theta_{13} \geq 0.025$ for ALL values of δ_{CP} .

More than an order of magnitude improvement over the current 90% CL

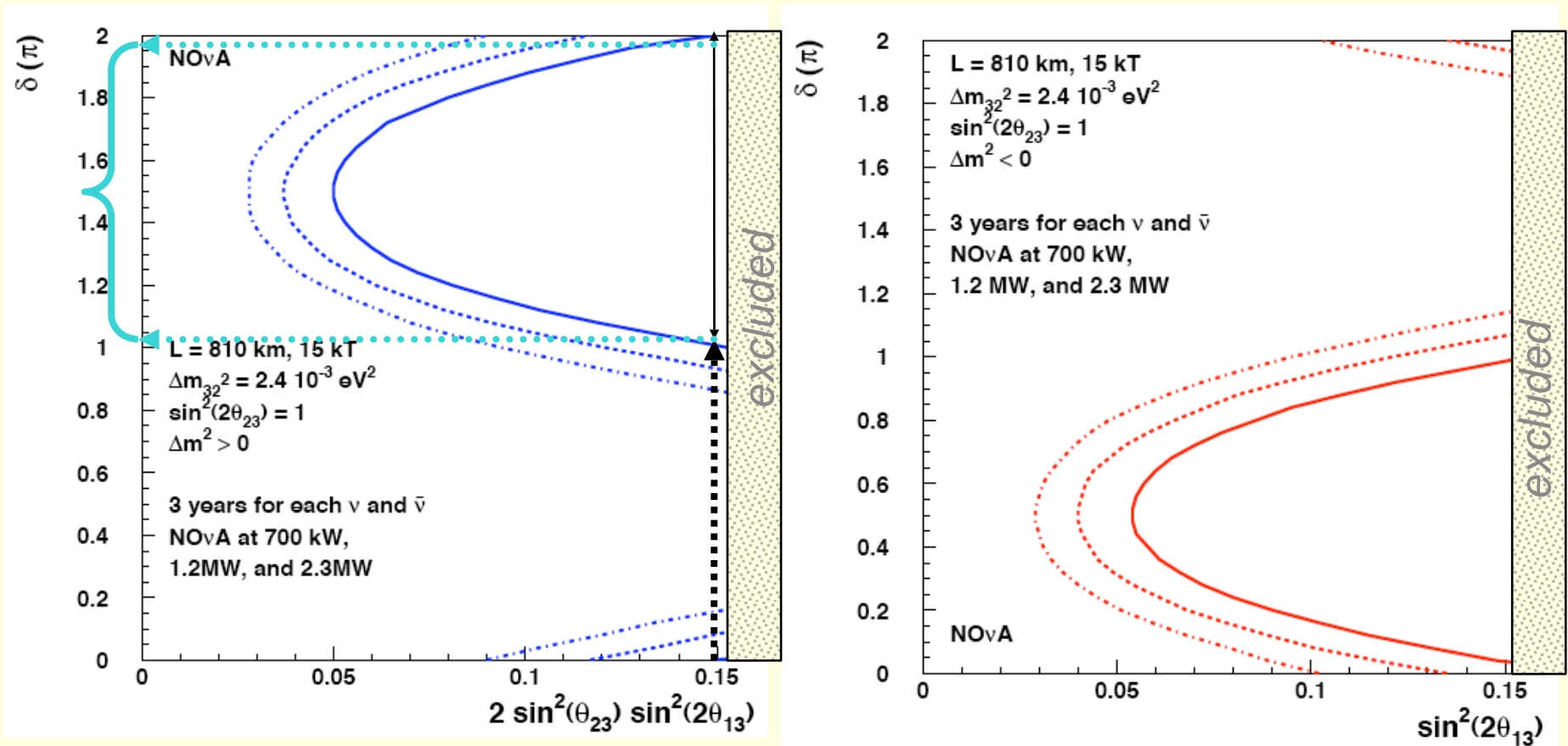


NOvA Sensitivity to the Mass Hierarchy



Interpreting NO ν A Sensitivity to the Mass Hierarchy

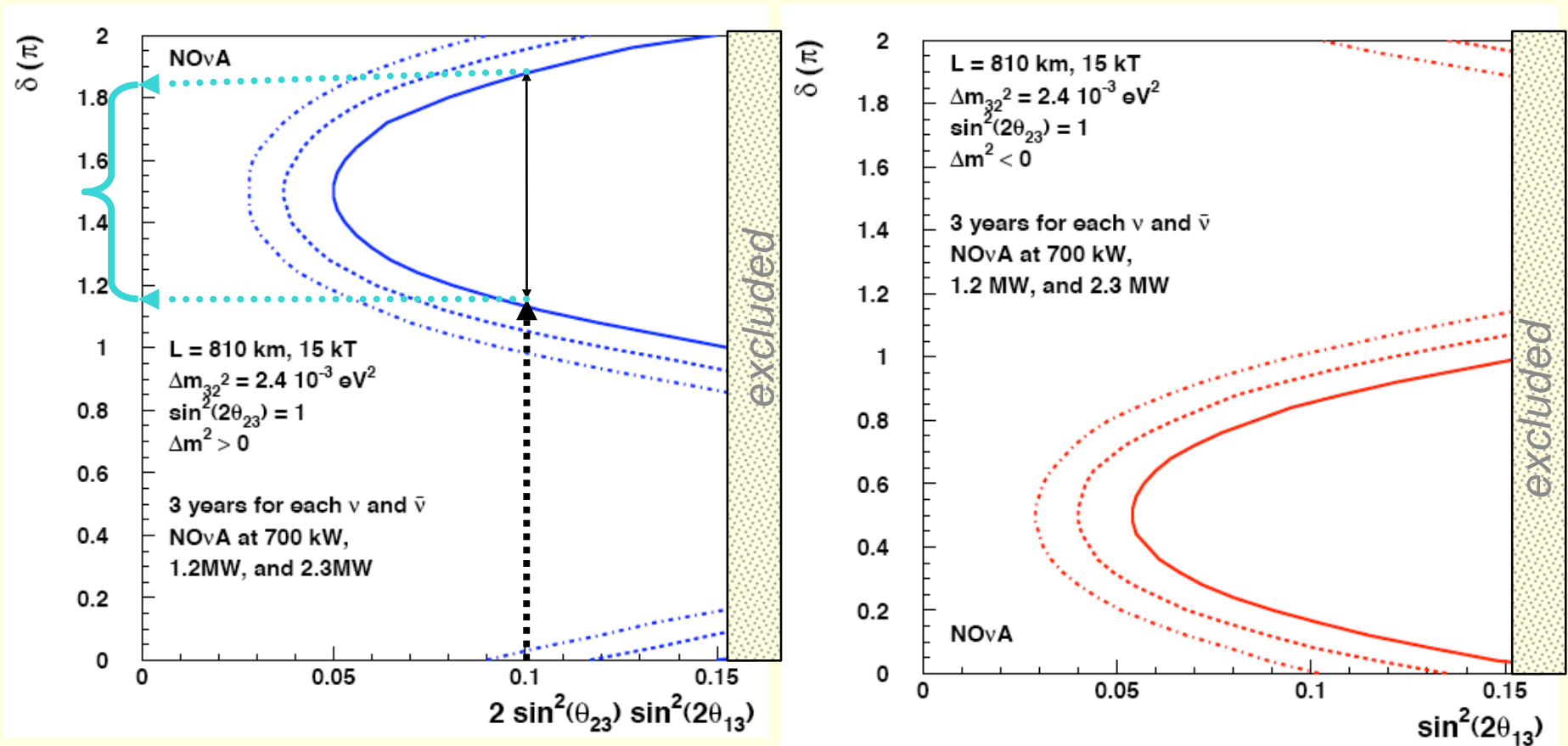
95% CL



If $\sin^2 2\theta_{13} = 0.15$, for 50% of the possible values of δ_{CP} the mass hierarchy can be determined at 95%CL

Interpreting NO ν A Sensitivity to the Mass Hierarchy

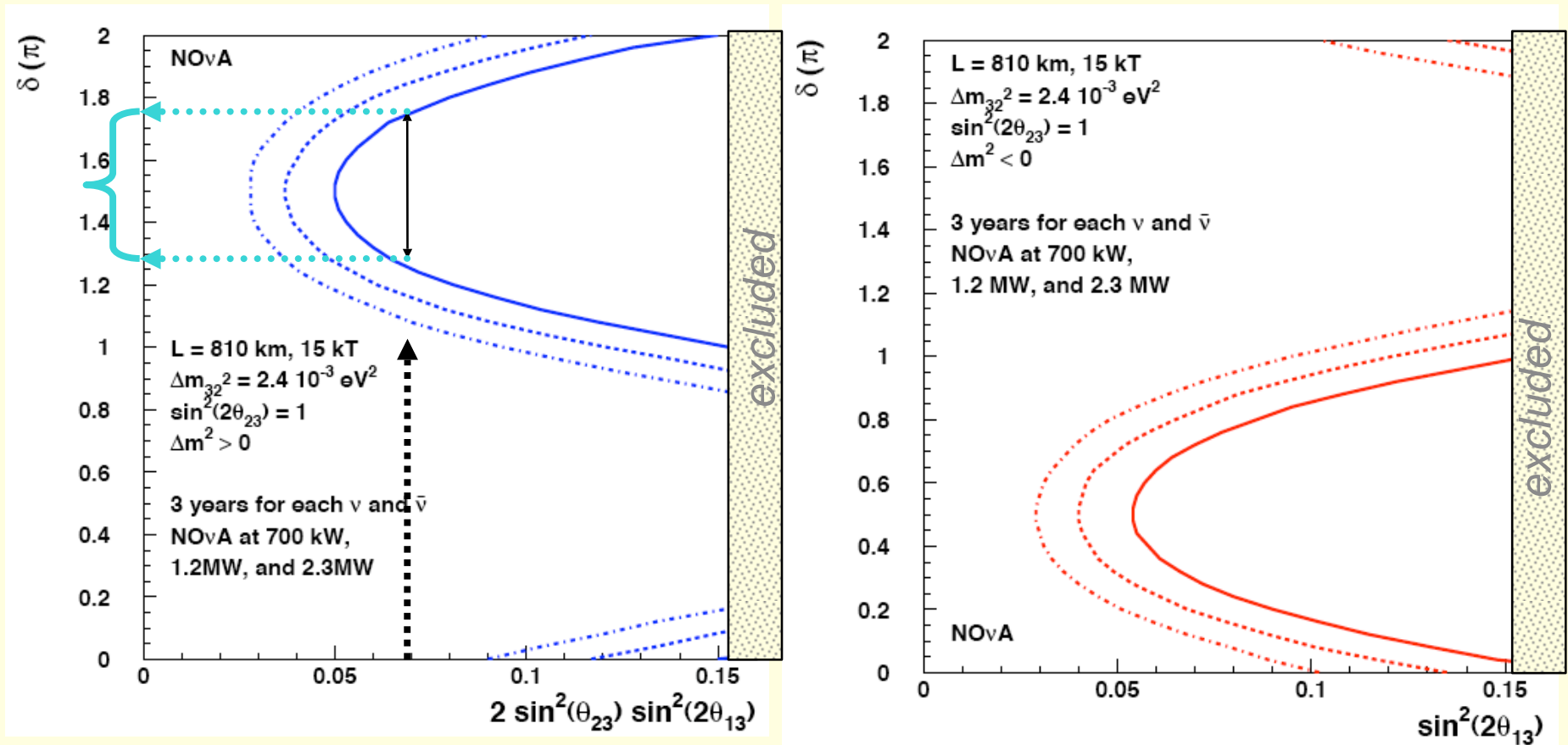
95% CL



If $\sin^2 2\theta_{13} = 0.10$, for 36% of the possible values of δ_{CP} the mass hierarchy can be determined at 95%CL

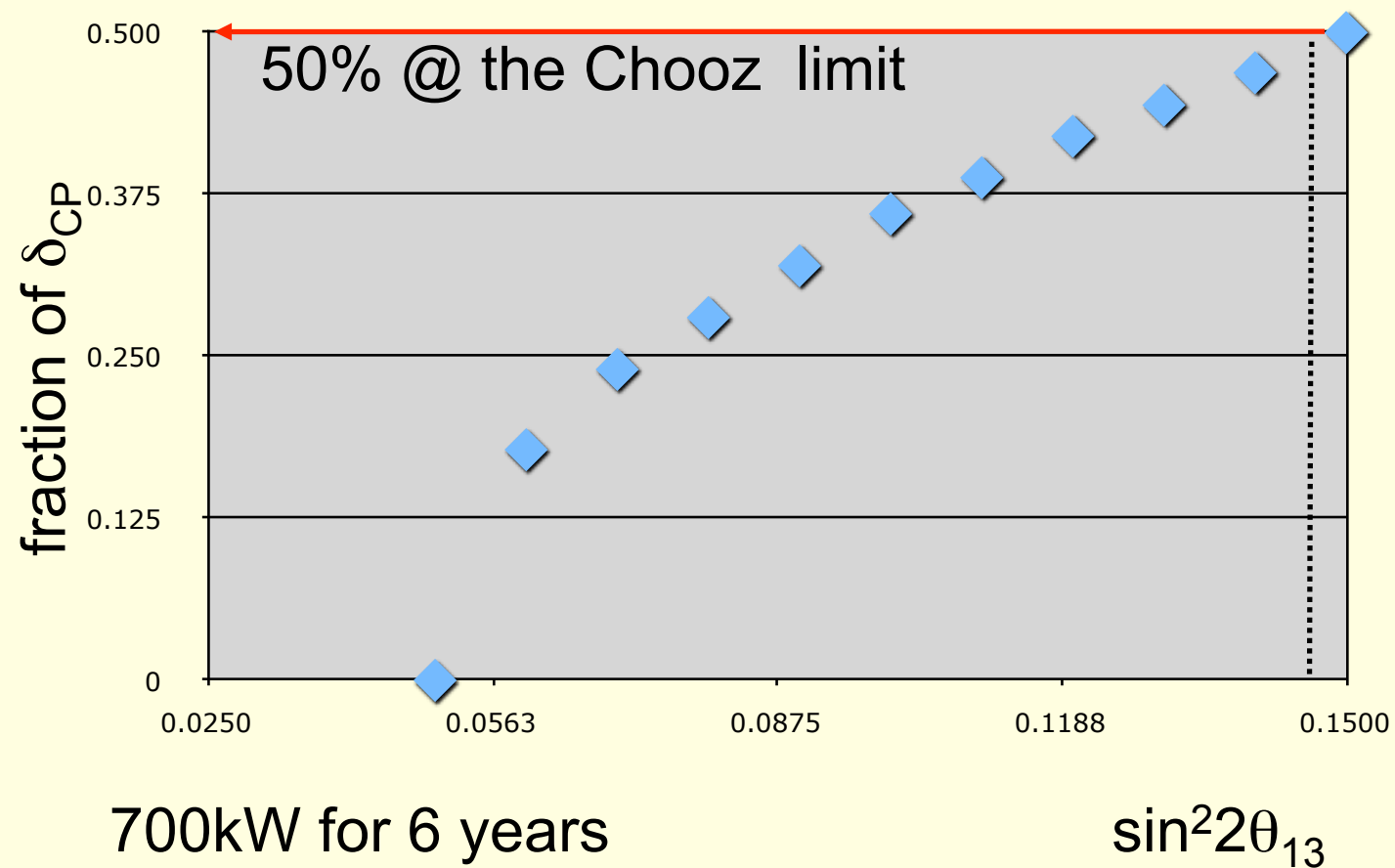
Interpreting NO ν A Sensitivity to the Mass Hierarchy

95% CL

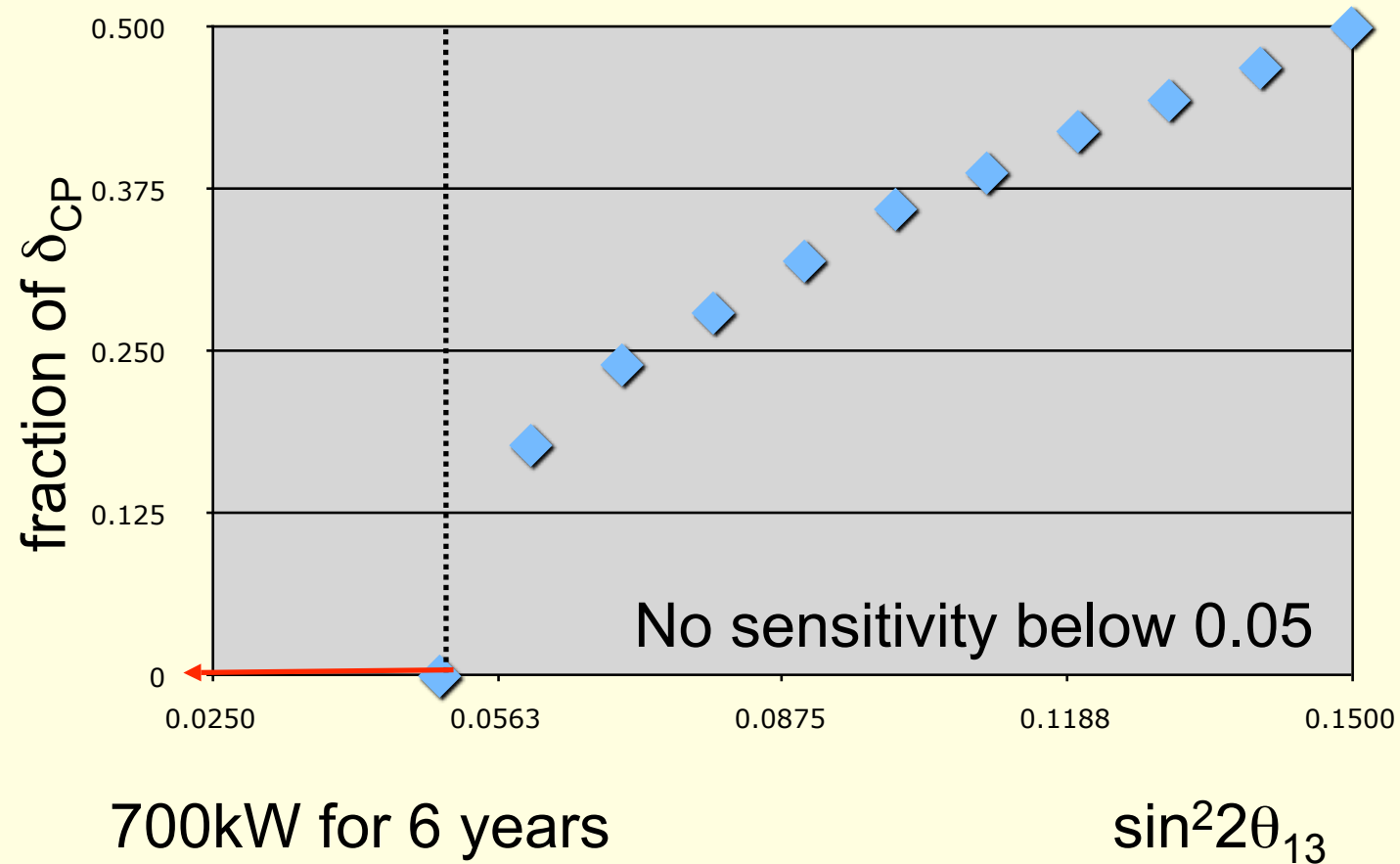


If $\sin^2 2\theta_{13} = 0.07$, for 24% of the possible values of δ_{CP} the mass hierarchy can be determined at 95%CL

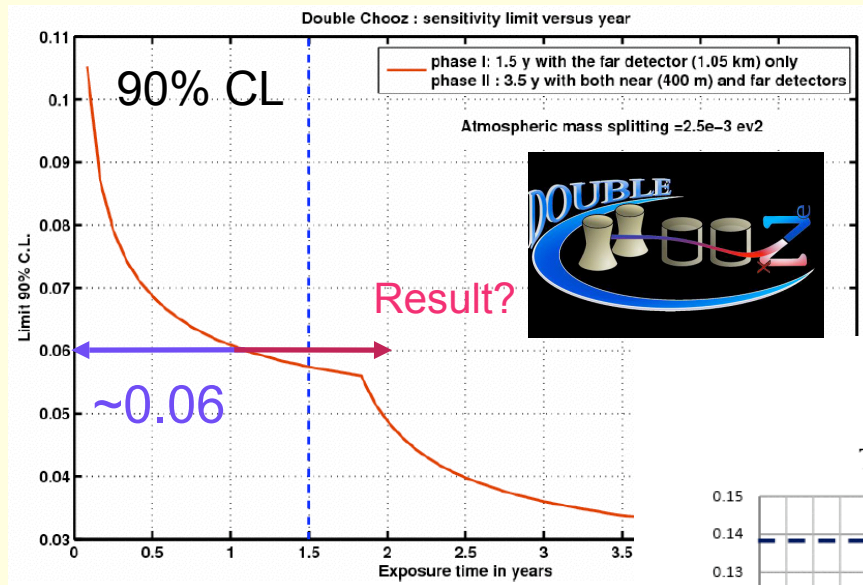
NO ν A 95% CL sensitivity to the Mass Hierarchy



NO ν A 95% CL sensitivity to the Mass Hierarchy

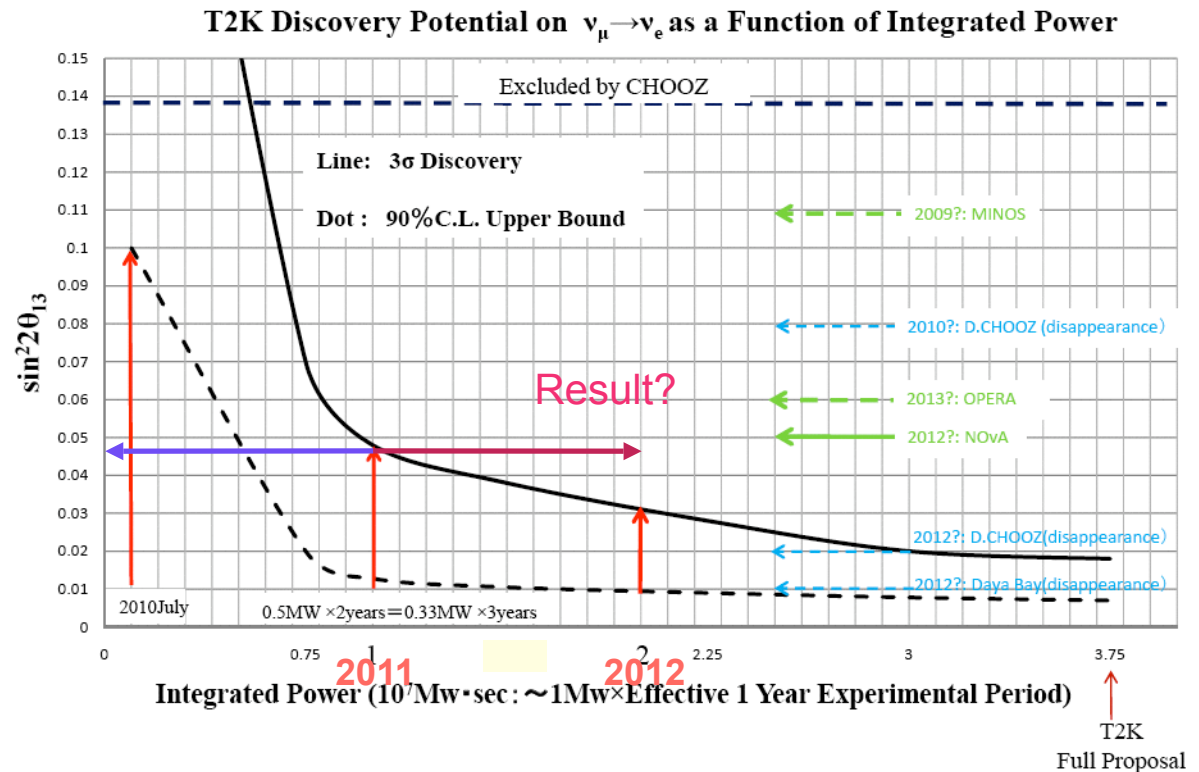
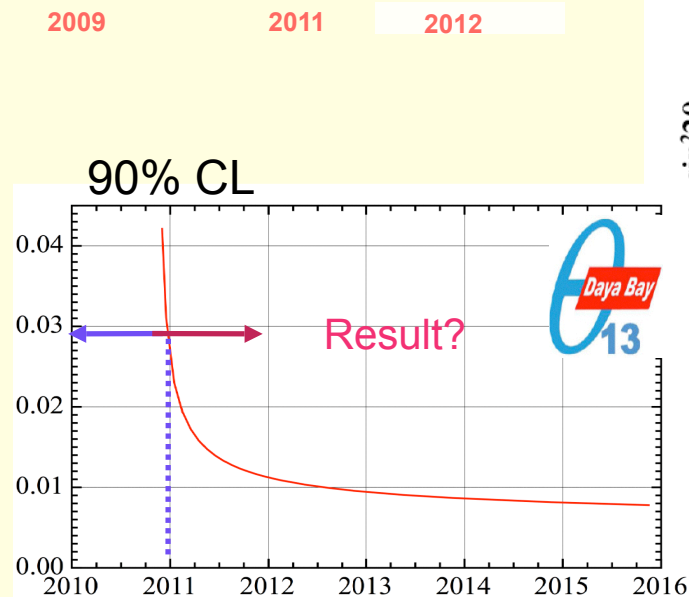


What are the prospects for knowing $\sin^2 2\theta_{13}$?



By 2012, we should have
a good indication if $\sin^2 2\theta_{13} > 0.05$

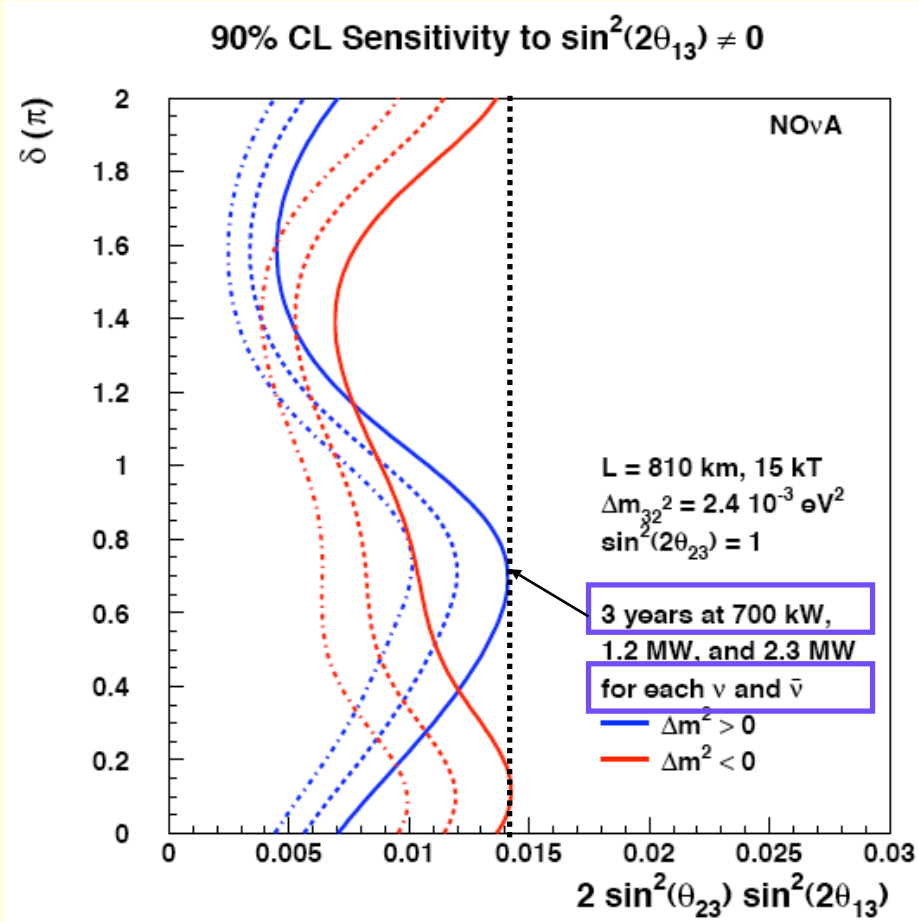
Takuya Hasegawa - NNN08



Neutrino Program Evolution beyond the “Phase I” θ_{13} experiments

- Numerous studies over the past several years have laid out options for further exploring the neutrino sector
 - In particular, searching for CP violation
- i.e. BNL-FNAL US long baseline neutrino experiment study (March 2006-June 2007) explored
 - Beam options
 - NuMI , new Wide Band Beam at a longer baseline
 - On and off axis detector locations
 - Detector technology options
 - Water cerenkov, liquid argon
- These studies make sense in the context of a non-zero determination of θ_{13}

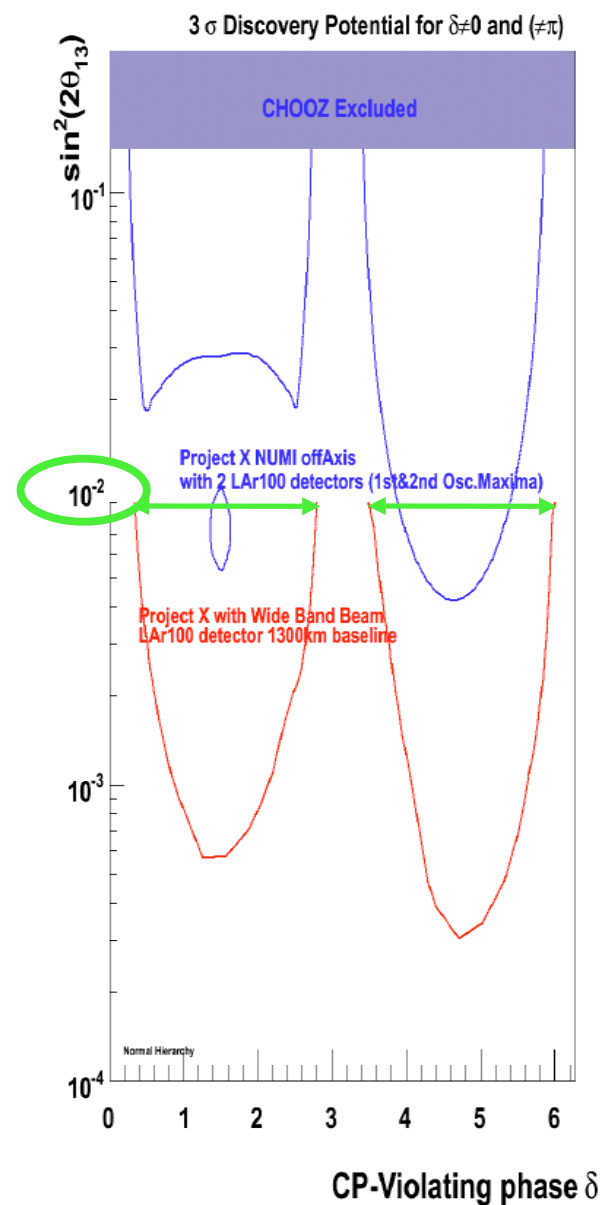
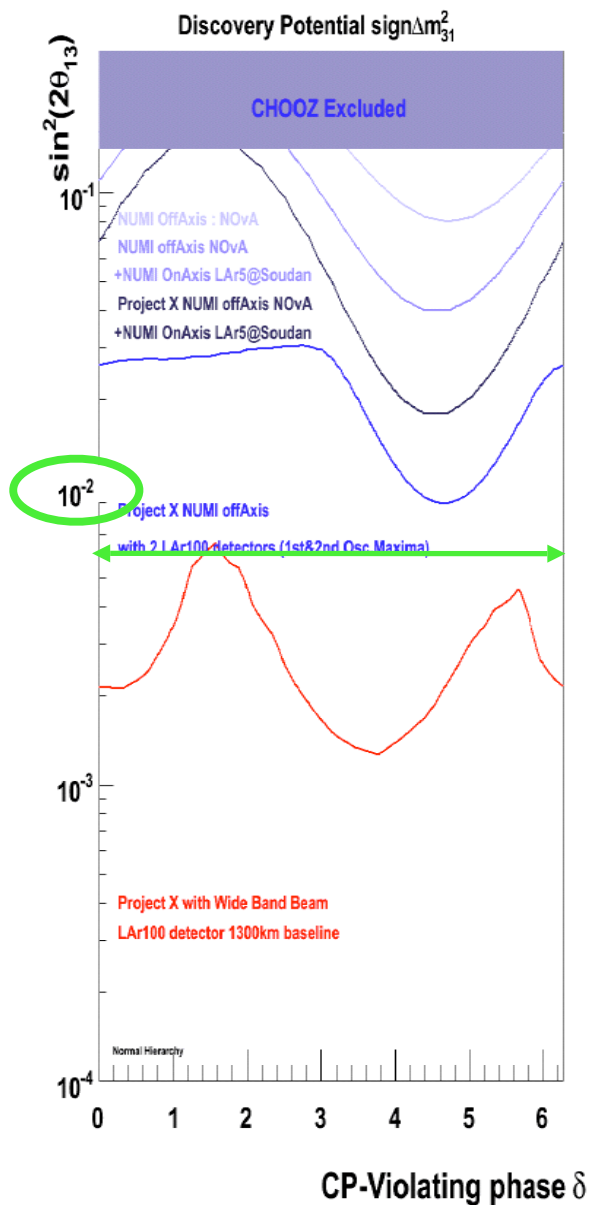
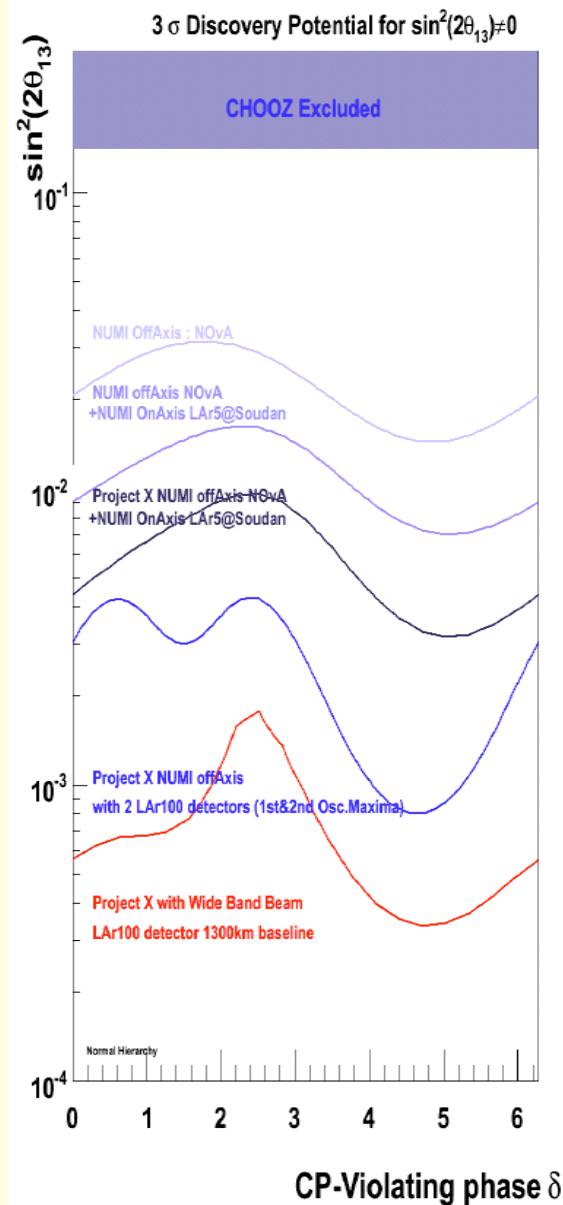
NO ν A Sensitivity for small $\sin^2 2\theta_{13}$



We can reach a 90% CL limit
 for $\sin^2 2\theta_{13} < 0.015$
 for ALL values of δ_{CP} .

General Conclusions about the next phase

- Future experiments using conventional* neutrino beams can be designed to have $3\text{-}5\sigma$ discovery potential for measuring **CP violation and the neutrino mass hierarchy** for values of $\sin^2 2\theta_{13}$ as low as ~ 0.01
- These sensitivities are reached assuming :
 - a **proton source** at the Megawatt level (or decades of running time)
 - a **neutrino beam** optimized to the oscillation probability (covering the 1st and 2nd oscillation maximum)
 - an **experiment baseline > 1000 km** (to improve the sensitivity to determine the mass hierarchy)
 - a **Detector** with effective mass (mass*efficiency) **$> 100\text{kT}$**
- *If nature has made θ_{13} very small we may need to consider a non-conventional neutrino source, i.e. **neutrino factory**



Plot by N. Saoulidou for Fermilab Steering Group

The Intensity Frontier

The accelerator-based neutrino program

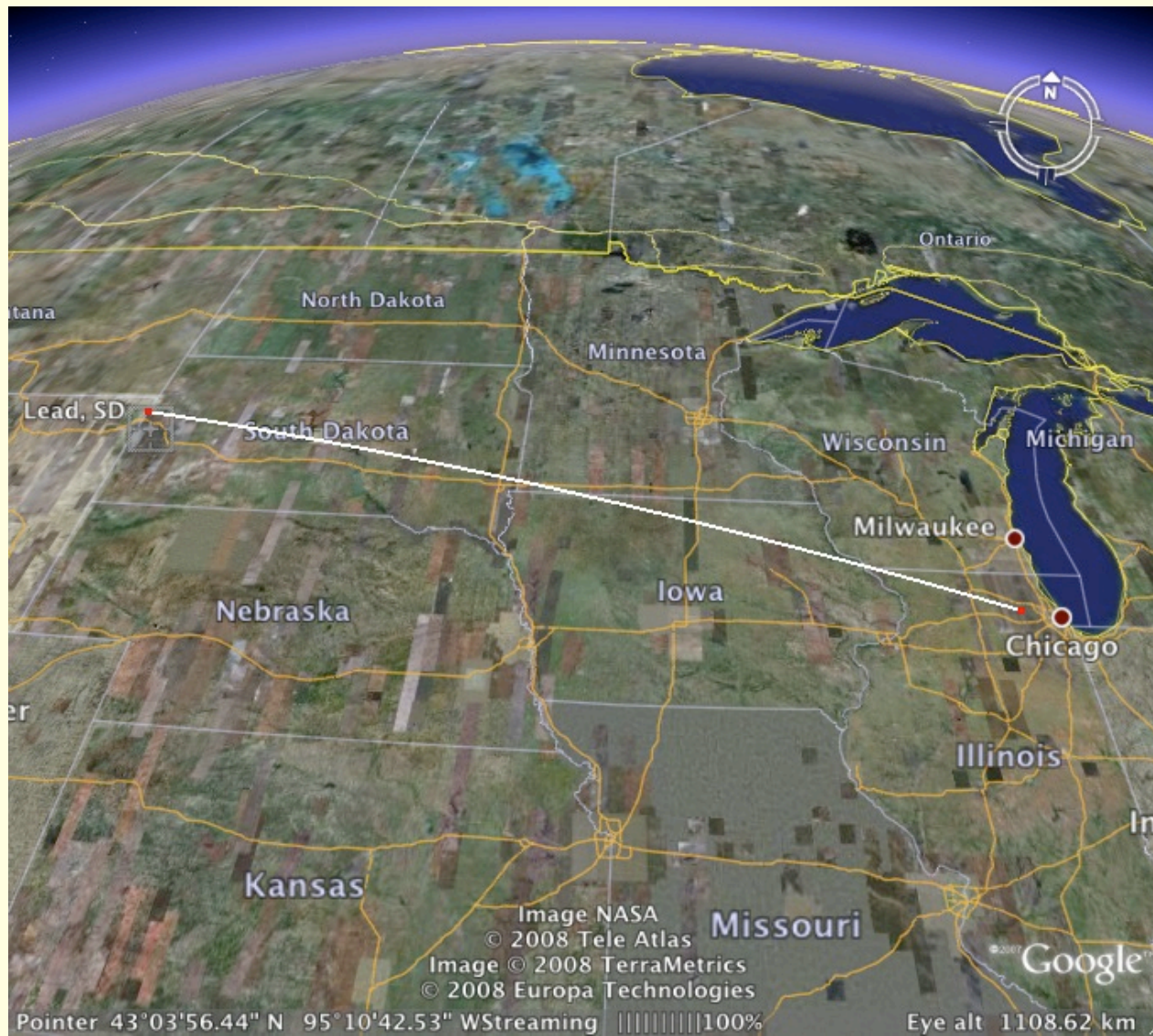
- Measurements of the mass and other properties of neutrinos are fundamental to understanding physics beyond the Standard Model and have profound consequences for understanding the evolution of the universe. The US can build on the unique capabilities and infrastructure at Fermilab, together with the proposed DUSEL, the Deep Underground Science and Engineering Laboratory proposed for the Homestake Mine, to develop a world-leading program in neutrino science. Such a program will require a multi-megawatt proton source at Fermilab.
- The panel recommends a world-class neutrino program as a core component of the US program, with the long-term vision of a large detector in the proposed DUSEL laboratory and a high-intensity neutrino source at Fermilab.

from P5 report

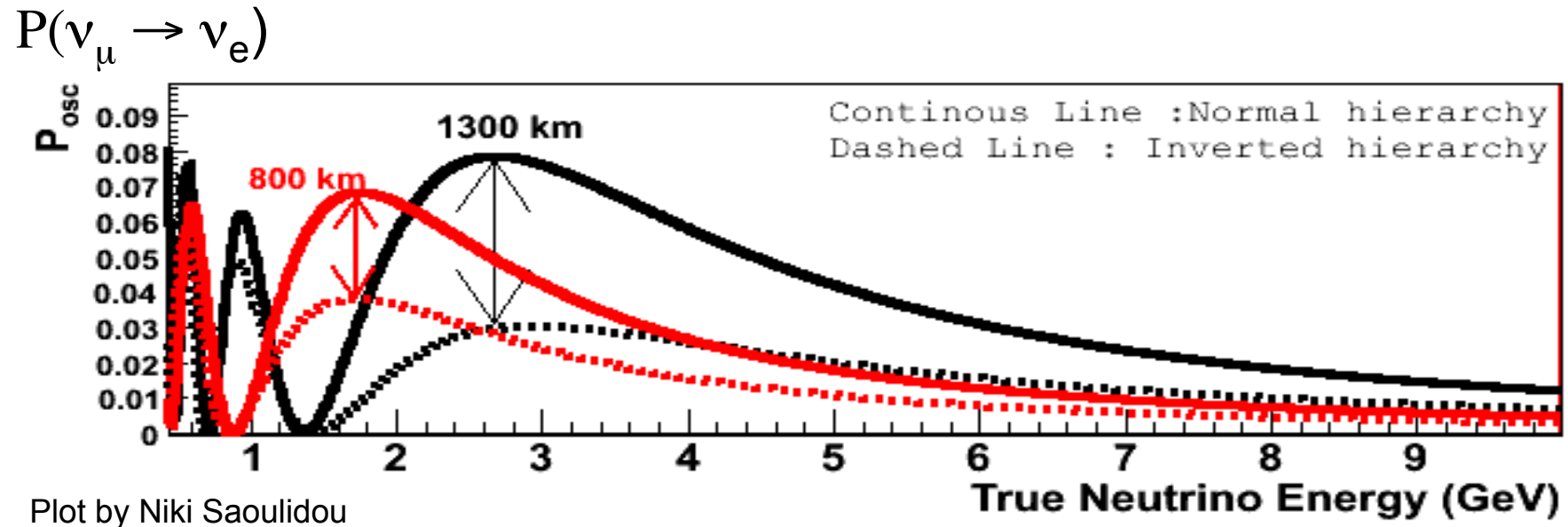
Neutrino Program (cont)

- The panel recommends proceeding now with an R&D program to design a multi-megawatt proton source at Fermilab and a neutrino beamline to DUSEL and recommends carrying out R&D on the technology for a large detector at DUSEL.
- Construction of these facilities could start within the period considered by this report.
- A neutrino program with a multi-megawatt proton source would be a stepping stone toward a future neutrino source, such as a neutrino factory based on a muon storage ring, if the science eventually requires a more powerful neutrino source. This in turn could position the US program to develop a muon collider as a long-term means to return to the energy frontier in the US

Fermilab to Homestake DUSEL (1290km)

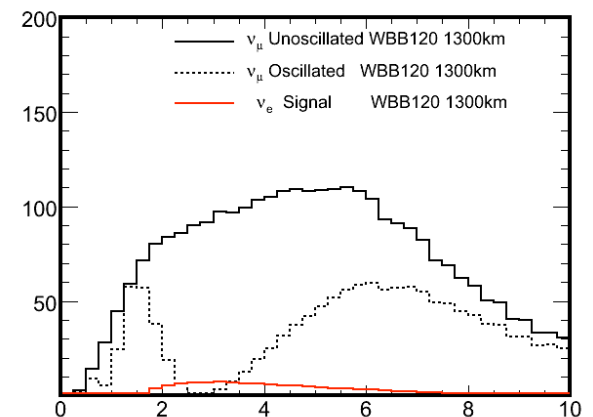
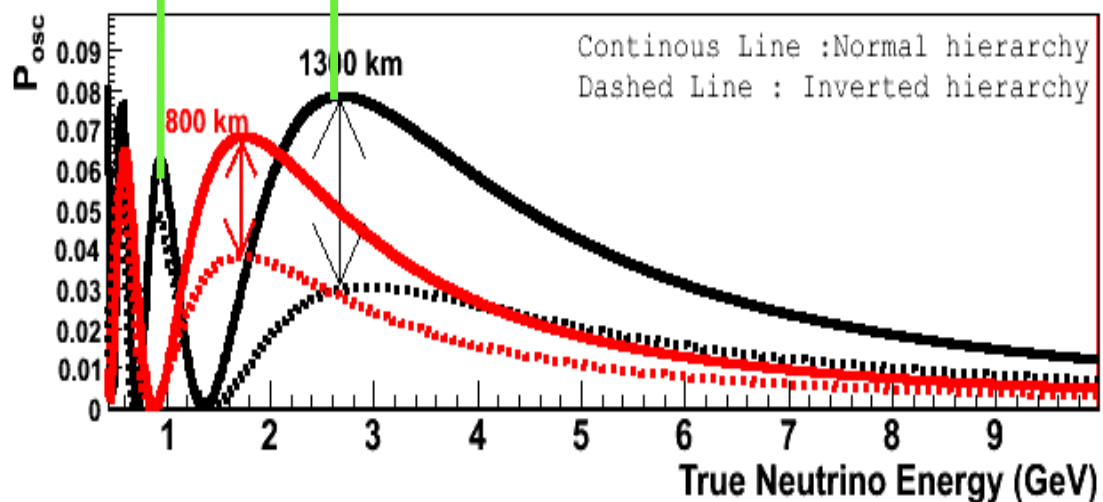
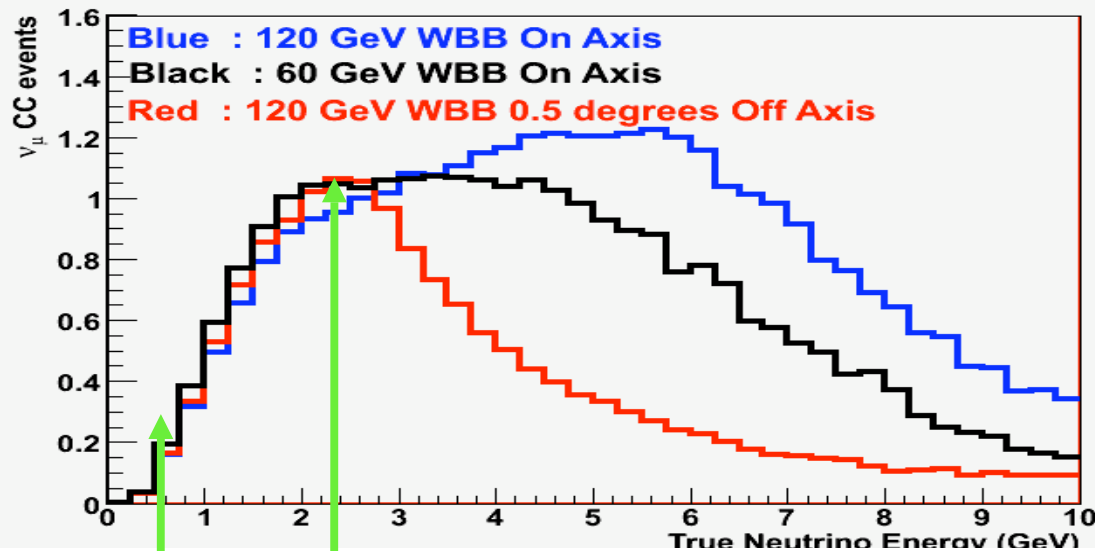


Advantage of the longer baseline



- Oscillation maxima are moved to higher energy
- Matter effects are significantly larger

The Experimental Technique : optimize the spectrum to the oscillation probability



		Neutrino Rates				Anti Neutrino Rates			
Beam (mass ordering)	$\sin^2 2\theta_{13}$	δ_{CP} deg.							
		0°	-90°	180°	+90°	0°	-90°	180°	+90°
NuMI LE 12 km offaxs (+)	0.02	76	108	69	36	20	7.7	17	30
NuMI LE 12 km offaxs (-)	0.02	46	77	52	21	28	14	28	42
NuMI LE 12 km offaxs (+)	0.1	336	408	320	248	86	57	78	106
NuMI LE 12 km offaxs (-)	0.1	210	280	224	153	125	95	126	157
NuMI LE 40 km offaxs (+)	0.02	5.7	8.8	5.1	2.2	2.5	1.6	0.7	3.3
NuMI LE 40 km offaxs (-)	0.02	4.2	8.0	5.7	2.0	2.3	2.2	0.8	3.6
NuMI LE 40 km offaxs (+)	0.1	17	24	15	9.4	6.7	2.8	4.6	8.5
NuMI LE 40 km offaxs (-)	0.1	12	21	16	7.7	6.6	3.4	6.4	9.6
WBLE 1300 km (+)	0.02	141	192	128	77	19	11	18	36
WBLE 1300 km (-)	0.02	58	111	88	35	45	25	45	64
WBLE 1300 km (+)	0.1	607	720	579	467	106	67	83	122
WBLE 1300 km (-)	0.1	269	388	335	216	196	154	196	240
WBLE 2500 km (+)	0.02	61	103	88	46	11	4.6	4.7	11
WBLE 2500 km (-)	0.02	16	36	33	13	28	15	18	31
WBLE 2500 km (+)	0.1	270	361	328	238	27	13	13	28
WBLE 2500 km (-)	0.1	47	92	85	39	103	74	80	109

Charge current
events per
100kT mass
per 1 MW per
10⁷ sec

No detector model
or backgrounds

(NuMI - 120 GeV
WBLE - 60 GeV)

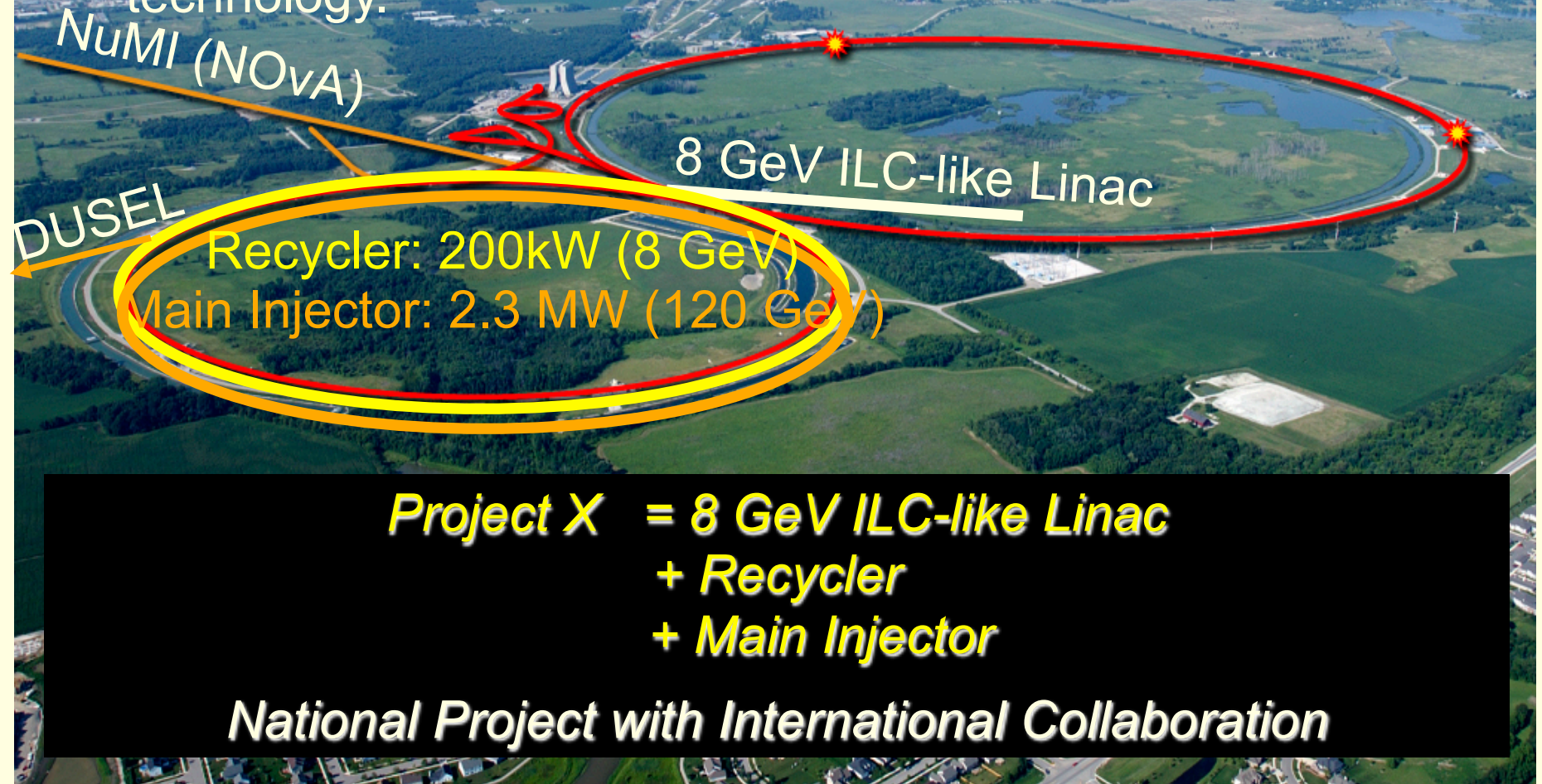
DUSEL
rates

~10-1000 evts

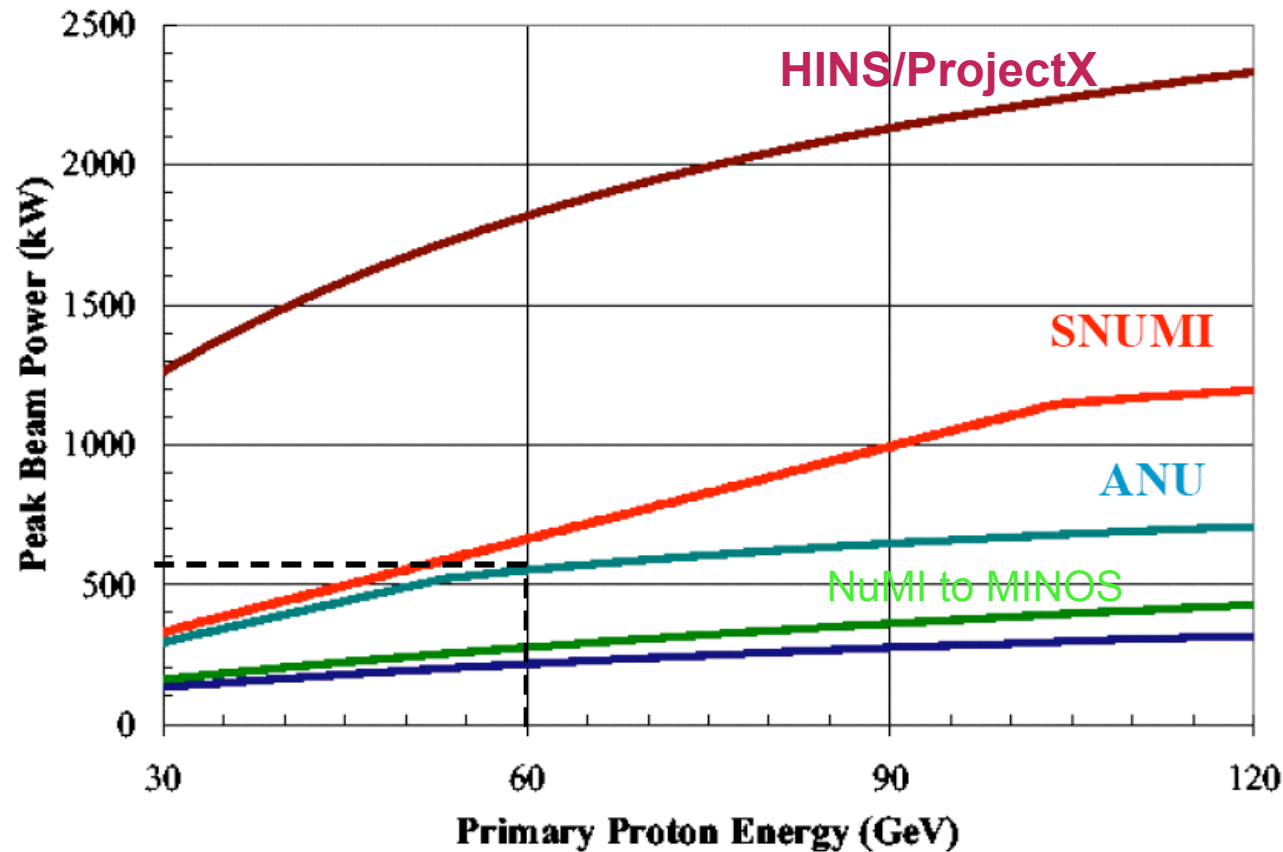
From BNL/FNAL study
(M. Bishai, B. Virin, M.
Dierkerson)

Fermilab vision :The Intensity Frontier with Project X:

Great flexibility toward a very high power facility while simultaneously advancing energy-frontier accelerator technology.



Plot courtesy : B. Zwaska



20x10²⁰ POT/yr

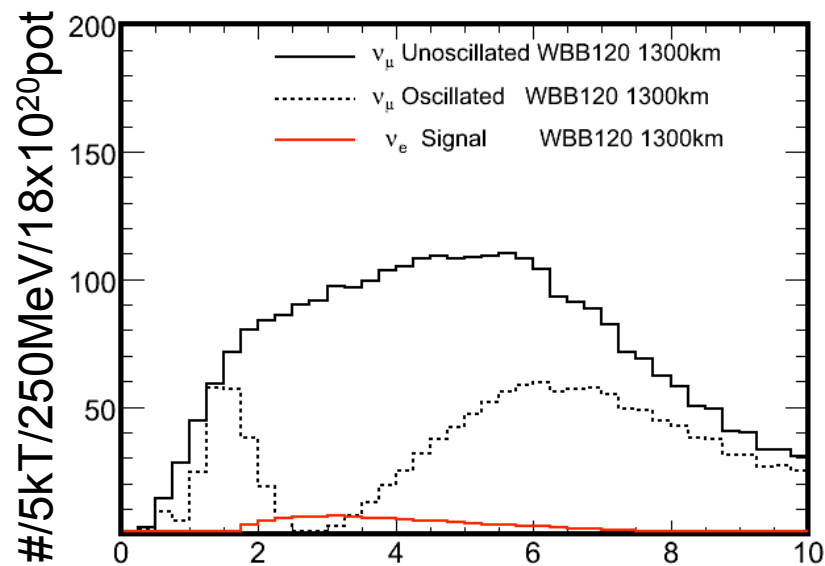
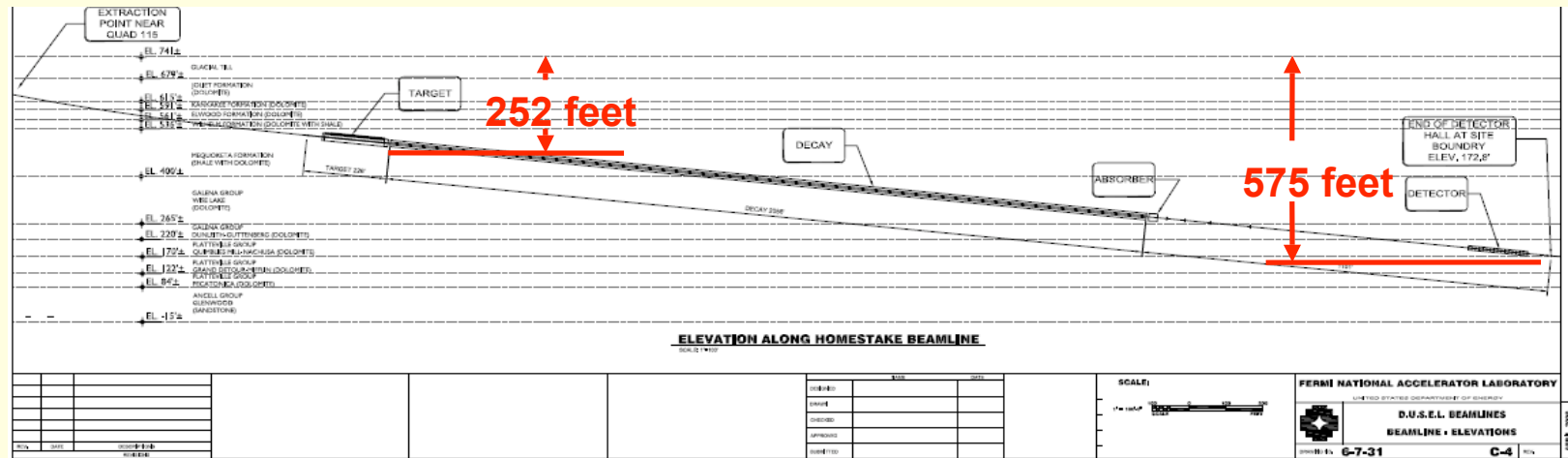
10x10²⁰ POT/yr

6x10²⁰ POT/yr

3x10²⁰ POT/yr

$$POT(10^{20}) = \frac{1000 \times BeamPower(MW) \times T(10^7 s)}{1.602 \times E_p(GeV)}$$

A beam to DUSEL : shorter & wider than NuMI

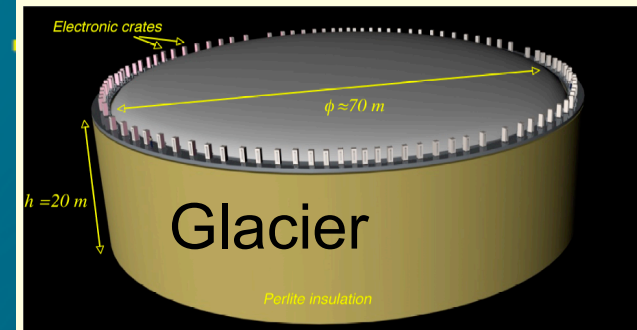
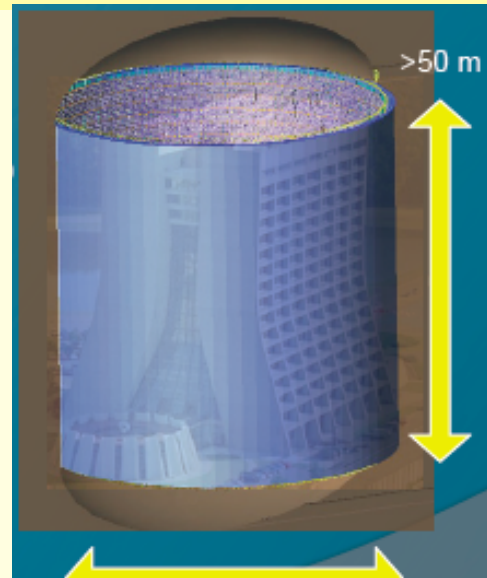
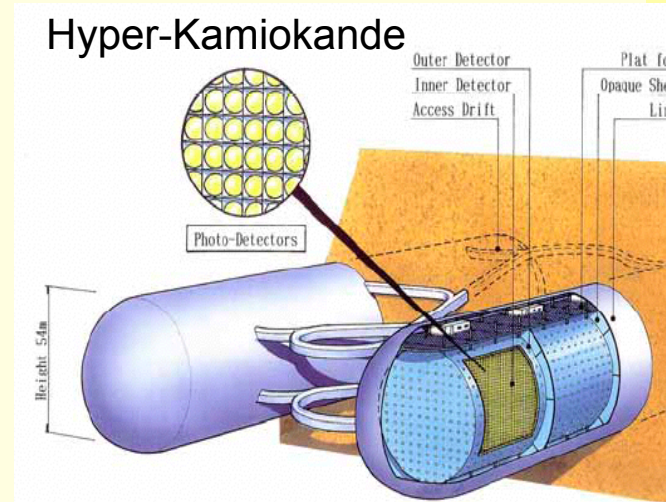
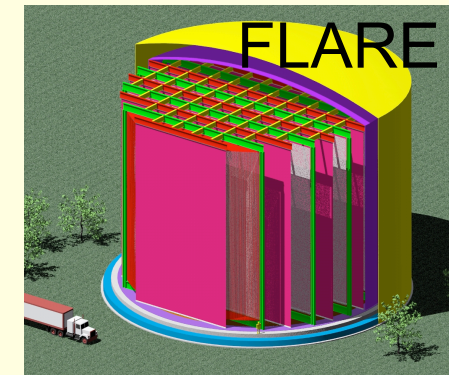


High power issues:
 groundwater activation,
 radioactive air emissions,
 target stress, radiation damage,
 decay pipe stress....

A **super beam** needs a **super detector**

World Wide Concepts for Large Detector

Water Cerenkov Liquid Argon Liquid Scintillator



DETECTOR LAYOUT

LENA

Cavern

height: 115 m, diameter: 50 m
shielding from cosmic rays: ~4,000 m.w

Muon Veto

plastic scintillator panels (on top)
Water Cherenkov Detector
1,500 phototubes
100 kt of water
reduction of fast neutron background

Steel Cylinder

height: 100 m, diameter: 30 m
70 kt of organic liquid
13,500 phototubes

Buffer

thickness: 2 m
non-scintillating organic liquid
shielding external radioactivity

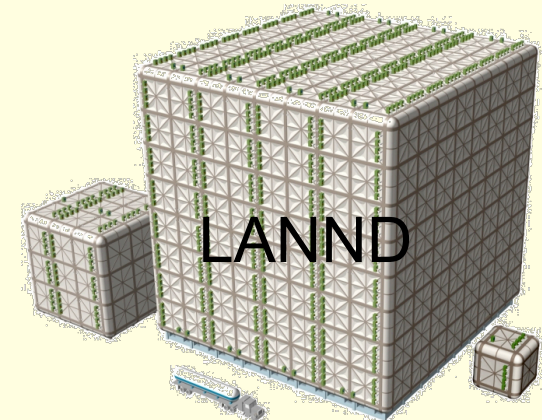
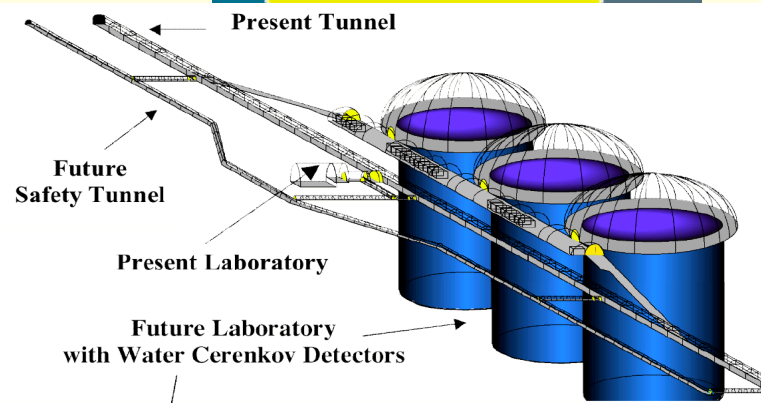
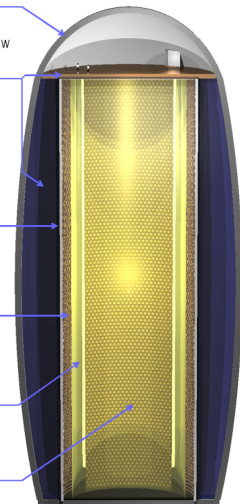
Nylon Vessel

parting buffer liquid from liquid scintillator

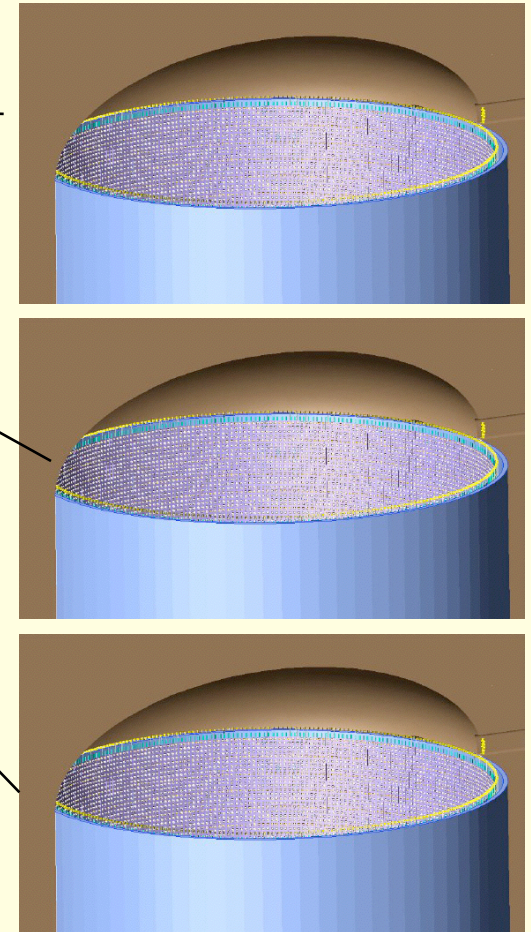
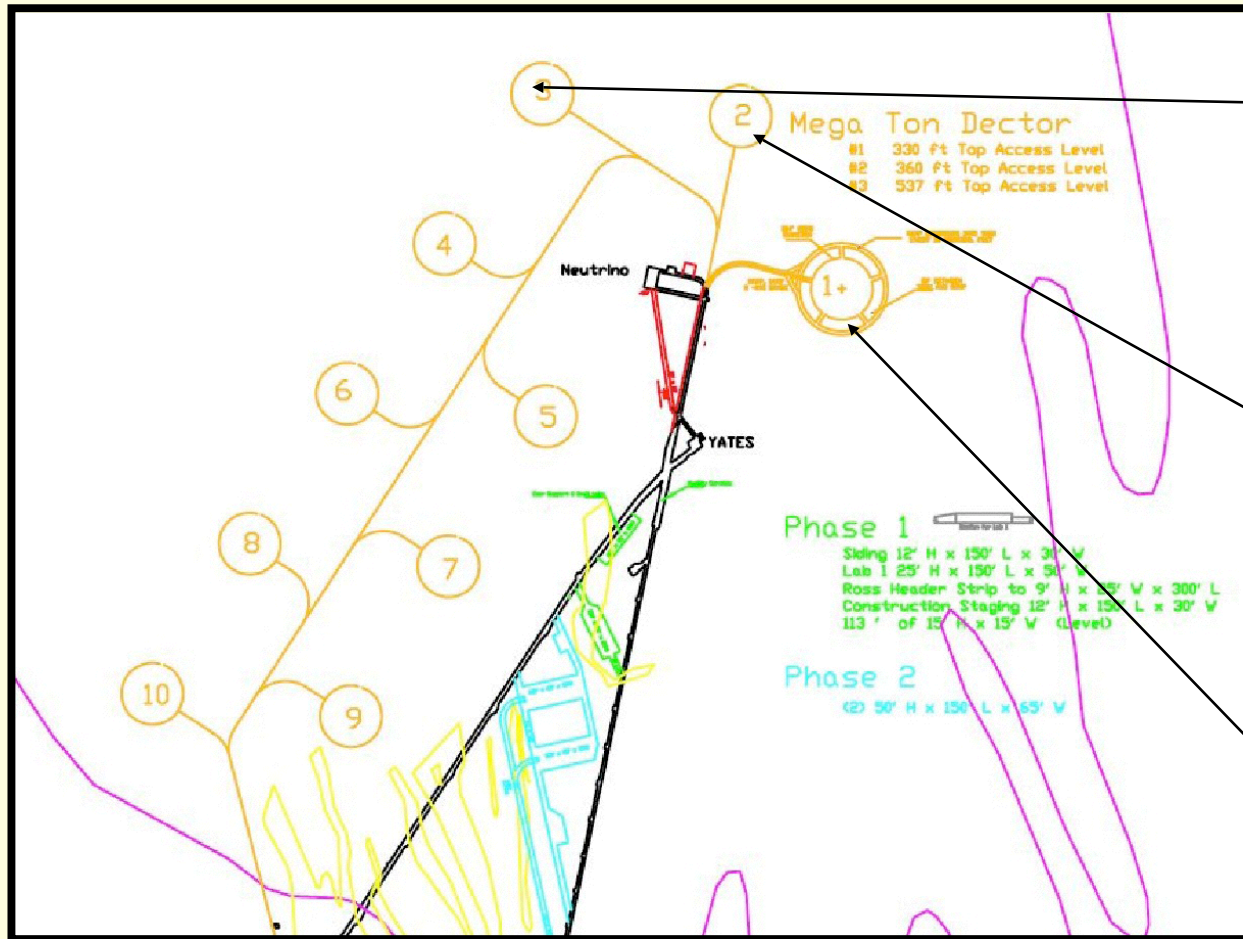
Target Volume

height: 100 m, diameter: 26 m
50 kt of liquid scintillator

vertical design is favourable in terms of rock pressure and buoyancy forces

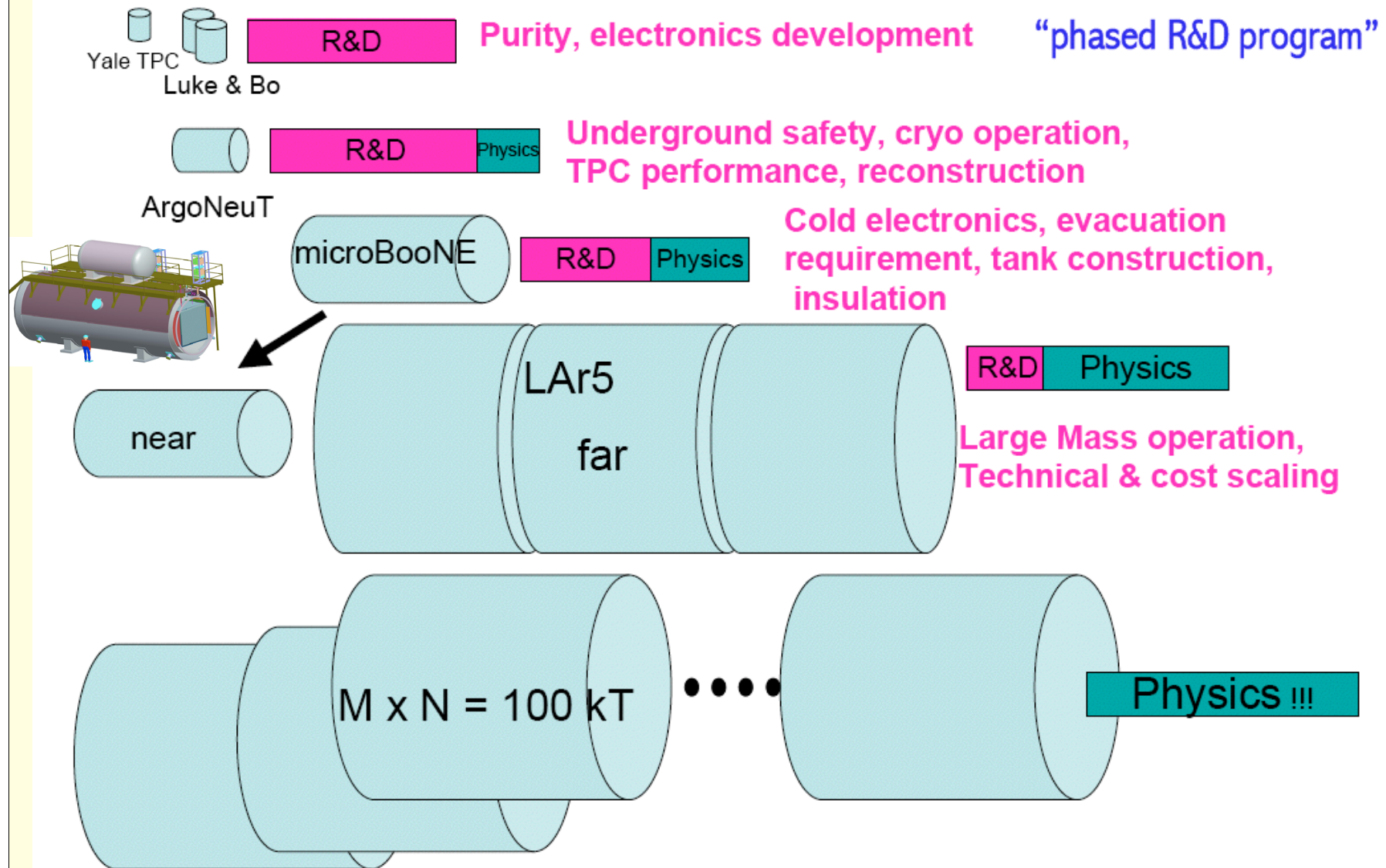


WC-100 x 3 @ Homestake DUSEL

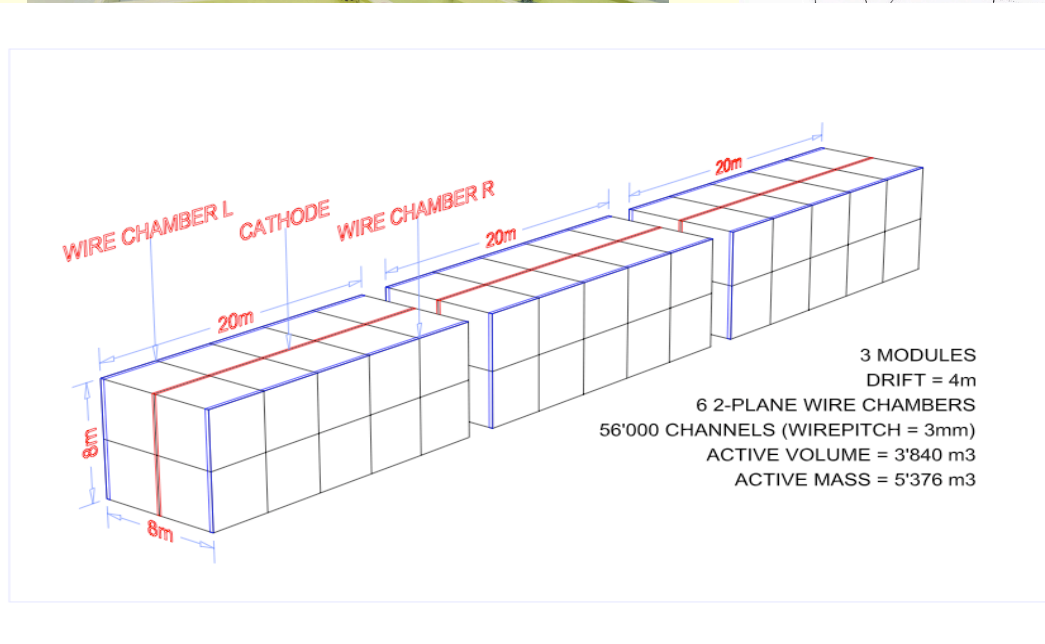
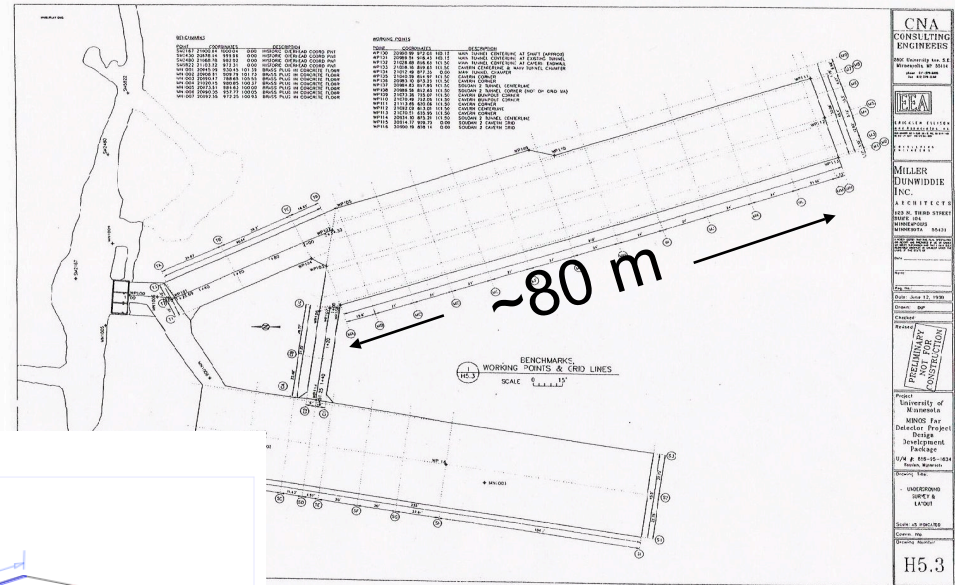
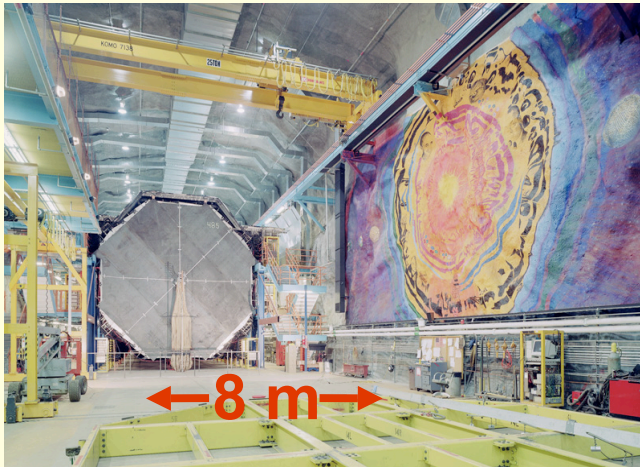


25% PMT coverage → 60,000 10 inch PMT's per module

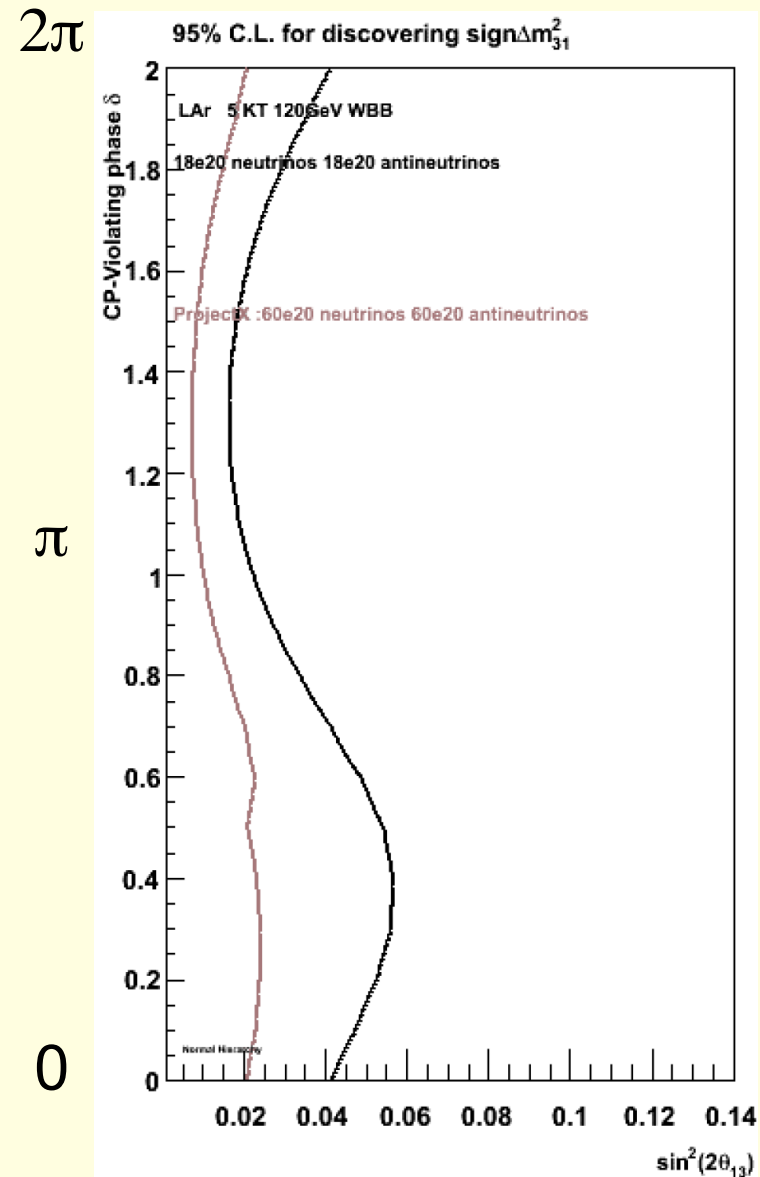
Evolution of the Liquid Argon Physics Program



Concepts for LAr5

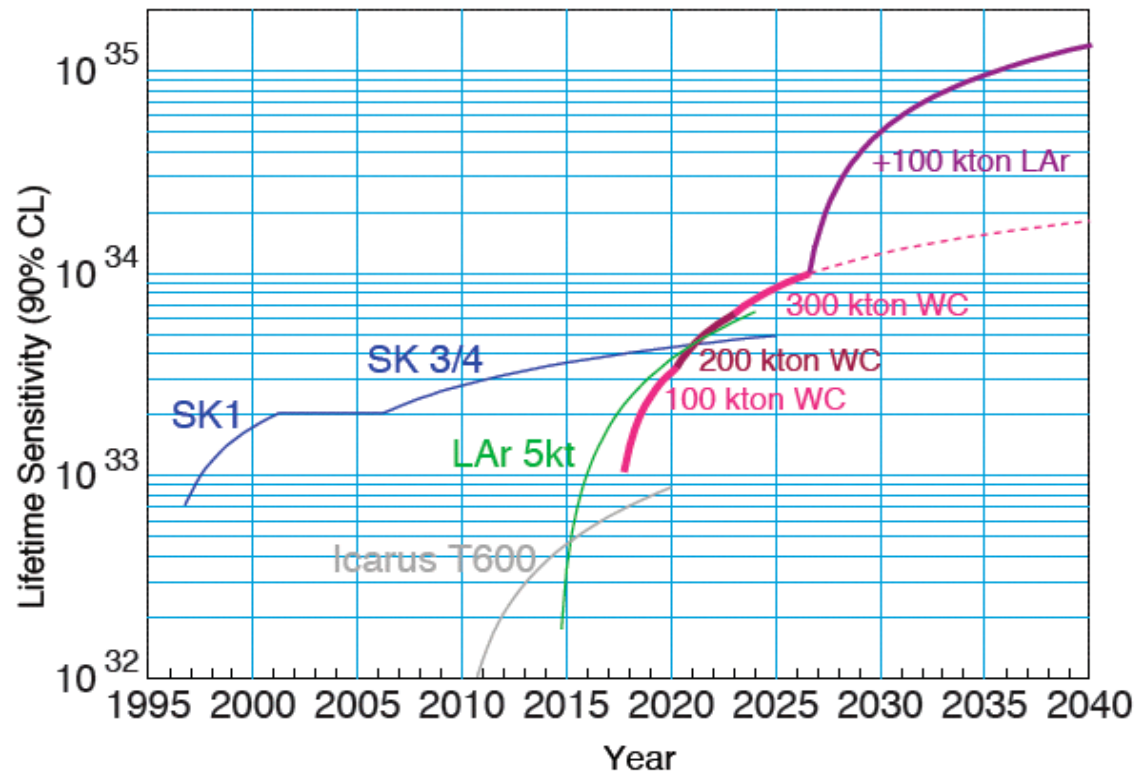


Physics reach of 5 kT LAr



at Homestake
DUSEL

An added bonus, while waiting for the new neutrino beam...



WC efficiency = 0.14
BG = 1.2 evts/100 kty
Nobs = Nbg

LAr efficiency = 0.98
BG = 0.1 evts/100 kty
Nobs = Nbg

Key DUSEL Dates

- July 2007: Homestake Site Selection for proposed DUSEL
- October 2008: S4 solicitation for experimental proposals
- December 2010: PDR-Preliminary Design Report, Baseline Scope, Schedule, Budget
- March 2011: Earliest National Science Board Presentation of DUSEL MREFC proposal
- October 2012: FDR Final Design Report for construction start in FY2013

LB DUSEL “collaboration” Organization

- Several workshops/meetings since April
 - June 20 at FNAL
 - August 14 at FNAL
 - October 14-15 at BNL
- Next meeting : January on West Coast
- Temporary Executive Committee formed
- Forming an Institutional Board of “interested groups”
- Detector technology groups are preparing Proposals for the NSF S4 solicitation
- To subscribe to the mailing list and get involved go to
<https://solid.physics.ucdavis.edu/mailman/listinfo/lbdusel>

Conclusions

- Over the past decade we have seen many exciting results from neutrino oscillation experiments looking at solar, atmospheric and accelerator neutrinos
 - We now know, to relatively good precision values for $\Delta m^2_{12}, \Delta m^2_{23}, \theta_{12}$ and θ_{23}
- Results from experiments to determine the third mixing angle, θ_{13} , are essential to laying out a strategy for further determination of the ν -mass-mixing matrix – in particular the parameter δ_{CP} , which will indicate whether or not CP is violated in the neutrino sector.

- If $\sin^2 2\theta_{13} \sim \geq 0.05$, with luck (and hard work) this result should be known by ~2012 from the Double Chooz, Daya Bay and T2K experiments
 - In this case, the NOvA experiment (which could/should start taking data in ~2013-14) will be able to confirm and contribute information about the mass hierarchy and δ_{CP}
 - Planning, leading to construction of a Phase II experiment, with a ν beam from Fermilab and massive detectors located at the DUSEL will offer the world wide neutrino community the opportunity to make precision measurements of neutrinos, as well as searches for proton decay and observation of astrophysical sources of neutrinos
 - A broad range of experiments at the DUSEL will make it a flagship facility for the Science community