# Total Neutrino and Antineutrino Nuclear Cross Sections around 1 GeV

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# Plan of the talk

#### 1 Introduction

present status of neutrino oscillations

what we know - we do not know future experimental facilities

- 2 Inclusive elastic neutrino-nucleus cross section
  - the IA regime
  - the  $(\nu_e^{-16}O)$  cross sections
  - comparison with other calculations
- 3 Resonance production
  - importance of including resonances in the calculation
  - total cross section
- 4 Conclusions and outlooks

# Introduction

• Present State of Neutrino Oscillation Parameters



STILL MISSING

 $\theta_{13}$  poorly known ( $\theta_{13} < 13^{\circ}$ )  $\longrightarrow$  3 Family Oscillations ?

 $\delta$  completely unknown  $\longrightarrow$  Leptonic CP-violation ?

High precision neutrino experiments required

to fully understand the neutrino mixing parameters

## Why cross sections around 1 GeV

many current and planned experiments use a u flux picked at  $\sim$  1 GeV

#### **MiniBoone**

T2K-I



and many others (NO $\nu$ A, high  $\gamma \beta$ -beams...)

- very few neutrino scattering data
- data have generically not been taken on the same targets used in the experiments

important to know very precisely the  $\nu$ -nucleus cross sections at  $E_{\nu} \sim 1 \text{ GeV}$ 

- at low energies ( $E_{\nu} \leq 0.6 0.7$  GeV): the dominant contribution comes from **quasi-elastic** scattering;
- at higher energies: inelastic production of charged leptons (via resonance excitation) + inelastic production of  $\pi^0$  also contribute
- negligible deep inelastic scattering contribution
- formalism to describe inclusive  $\nu + A \rightarrow l + X$  reaction



the problem is the calculation of the hadronic tensor  $W^{\mu
u}_A$ 

- for  $|\mathbf{q}| < 0.5 \,\mathrm{GeV}$  NMBT + nonrelativistic wave functions + expansion of the current operator in powers of  $|\mathbf{q}|/m_N$ Carlson&Schiavilla, Rev. Mod. Phys. 70, 743 (1998)
- for larger  $|\mathbf{q}|$  (to which we are interested in) we can no longer describe the final states  $|X\rangle$  in terms of nonrelativistic nucleons

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we need a set of simplifying assumptions to describe relativistic motion of final state particles and the occurrence of inelastic processes



Benhar et al.,

Phys.Rev.D72:053005,2005



$$W_{A}^{\mu\nu} = \frac{G_{F}^{2} V_{ud}^{2}}{2} \int d^{3}p \, dE \, P(\mathbf{p}, E) \frac{1}{4 E_{\mathbf{p}} E_{|\mathbf{p}+\mathbf{q}|}} W^{\mu\nu}(\tilde{p}, \tilde{q}) \, \delta(\nu - E - E_{|\mathbf{p}+\mathbf{q}|})$$

•  $P(\mathbf{p}, E)$  is the spectral function : probability distribution of finding a nucleon with momentum  $\mathbf{p}$  and removal energy E in the target nucleus

#### it encodes all the informations about the initial struck particle

- $W^{\mu
  u}$  describes electroweak interactions of the i-th nucleon in free space
- effect of nuclear binding of the struck nucleon is accounted for by the replacement  $q = (\nu, \mathbf{q}) \rightarrow \tilde{q} = (\tilde{\nu}, \mathbf{q})$  with  $\tilde{\nu} = E_{|\mathbf{p}+\mathbf{q}|} E_{\mathbf{p}}$

(a fraction  $\nu - \tilde{\nu}$  goes into excitation energy of the spectator system)

then we get

$$\frac{d^2 \sigma_{IA}}{d\Omega dE_l} = \int d^3 p \, dE \, P(\mathbf{p}, E) \, \frac{d^2 \sigma_{\text{elem}}}{d\Omega dE_l} \, \delta(\nu - E - E_{|\mathbf{p}+\mathbf{q}|})$$

$$\frac{d^2 \sigma_{\text{elem}}}{d\Omega dE_l} = \frac{G_F^2 V_{ud}^2}{32 \pi^2} \frac{|k'|}{|k|} \frac{1}{4 E_{\mathbf{p}} E_{|\mathbf{p}+\mathbf{q}|}} L_{\mu\nu} W^{\mu\nu}$$

• The hadronic tensor is decomposed in structure functions as usual

$$\begin{split} W^{\mu\nu} &= -g^{\mu\nu} W_1 + \tilde{p}^{\mu} \, \tilde{p}^{\nu} \, \frac{W_2}{m_N^2} + i \, \varepsilon_{\mu\nu\alpha\beta} \, \tilde{q}^{\alpha} \, \tilde{p}^{\beta} \, \frac{W_3}{m_N^2} + \tilde{q}^{\mu} \, \tilde{q}^{\nu} \, \frac{W_4}{m_N^2} + \\ (\tilde{p}^{\mu} \, \tilde{q}^{\nu} + \tilde{p}^{\nu} \, \tilde{q}^{\mu}) \, \frac{W_5}{m_N^2} \end{split}$$

• the formalism can be applied to **both** elastic and anelastic processes specifying the form of the structure functions  $W_i$ 

1- tensor contraction

 $A_1 = m_N^2 \left( k \cdot k' \right)$ 

$$L^{\mu\nu} W_{\mu\nu} = 16 \sum_{i} W_i \left(\frac{A_i}{m_N^2}\right)$$

 $A_{1} = (V_{N} \in V_{N}) = (K + \tilde{p}) = (K + \tilde{q}) = (K$ 



- 2- the spectral function  $P(\mathbf{p}, E)$
- we analyzed different models for  $P({\bf p},E)$  for  $^{16}O(\nu_e,e^-)$

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to start: the elementary cross section



quasi-elastic inclusive cross section

- 2- the spectral function  $P(\mathbf{p}, E)$ 
  - we analyzed different models for  $P(\mathbf{p},E)$  for  ${}^{16}O(\nu_e,e^-)$

a simple prescription: the Fermi gas

$$P(\mathbf{p}, E) = \left(\frac{6 \pi^2 A}{p_F^3}\right) \,\theta(p_F - \mathbf{p}) \,\delta(E_{\mathbf{p}} - \varepsilon + E)$$

where  $p_{\it F}$  is the Fermi momentum and  $\varepsilon$  is the average binding energy



- 2- the spectral function  $P(\mathbf{p}, E)$ 
  - we analyzed different model for  $P(\mathbf{p},E)$  for  ${}^{16}O(\nu_e,e^-)$

realistic spectral function

Benhar et al.,

Nucl. Phys. A 579 (1994) 493 Phys. Rev D72 (2005) 053005

- overwhelming evidence from electron scattering that the energy-momentum distribution of nucleons in the nucleus is quite different from that predicted by Fermi gas
- the most important feature is the presence of strong nucleon-nucleon (NN) correlations (virtual scattering processes leading to the excitation of the participating nucleons to states of energy larger than the Fermi energy)



#### spectral function extends to $|\mathbf{p}| \gg p_F$ and $E \gg \varepsilon$

- 2- the spectral function  $P(\mathbf{p}, E)$
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realistic spectral function Benhar et al., Nucl. Phys. A 579 (1994) 493



 realistic description must take into account interaction of struck nucleon with the spectator system

introduction of Final State Interaction (FSI)

• a simple prescription to include statistical correlation: Pauli blocking

 $P(\mathbf{p}, E) \rightarrow P(\mathbf{p}, E) \,\theta(|\mathbf{p} + \mathbf{q}| - \bar{p}_F)$ 

$$\bar{p}_F$$
 = average nuclear Fermi momentum =  $\int d^3 r \, \rho_A(\mathbf{r}) \, p_F(\mathbf{r}) = 209 \, MeV$   
 $p_F(\mathbf{r}) = \left[ 3 \, \pi^2 \, \rho_A(\mathbf{r}) / 2 \right]^{1/3} \quad \rho_A(\mathbf{r}) = \text{nuclear density}$ 



 Pauli blocking reduces the available phase space for knocked nucleon

differences between Fermi gas (with Pauli blocking) and our (red) prediction ranges from 15% to 4%

#### Elastic interactions Comparison with other calculations

 most of them in the low (and very low) neutrino energy region
 J.E. Amaro et al, Phys.Rev.C70:055503,2004, Erratum-ibid.C72:019902,2005: collection of cross-sections on <sup>16</sup>O including FSI, RPA (using an effective Nucleon-Nucleon force) and Coulomb distortion



good agreement with cross sections including SF + Pauli blocking the Fermi gas model is unsatisfactory

## **About Final State Interactions**

In quasi-elastic inclusive processes dynamical FSI are weak:

- an energy shift of the differential cross section, due to the fact that the struck nucleon feels the mean field generated by the spectator particles
- a redistribution of the strength, leading to the quenching of the peak and the enhancement of the tails





- solid line: Impulse Approximation + FSI
- dotted line: Fermi Gas

Benhar et al., Phys. Rev D72 (2005) 053005

FSI <u>do not</u> affect the <u>total</u> inclusive cross section, resulting from integration over the lepton variables

∜

• we are interested in the reactions  $\nu \ n \to l^- \ \Delta^+$  and  $\nu \ p \to l^- \ \Delta^{++}$ 

we use the <u>same</u> formalism as above where:

- the coefficients  $A_i$  are the same
- structure functions and form factors (magnetic dominance approximations) for  $\langle \Delta^{++} | J_{\mu} | p \rangle$  are taken from Lalakulich et al., Phys Rev. D 71, 074003 (2005)
- we used isospin relation to obtain  $\langle \Delta^{++} | J_{\mu} | p \rangle = \sqrt{3} \langle \Delta^{+} | J_{\mu} | n \rangle$



 strong reduction with respect to both the elementary and RFGM calculations

at 
$$E_{\nu} = 1 \text{ GeV}$$
  
 $\sigma_{SF} / \sigma_{RFGM} \sim 0.8$ 

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• contribution from  $\Delta$  production turns out to be important for  $E_{\nu} > 0.5~{\rm GeV}$ 

- we also evaluate the impact on the cross section of including higher resonances Lalakulich et al., Phys.Rev. D74,014009 (2005)
- three isospin 1/2 states:  $P_{11}(1440), D_{13}(1520), S_{11}(1535)$



resonance production

- as expected at the energies under discussion the second resonance region can be neglected
- $\sigma_{\Delta}: \sigma_{S_{11}}: \sigma_{D_{13}}: \sigma_{P_{11}} = 1: 0.12: 0.06: 0.02 \text{ at } E = 1.5 \text{ GeV}$

#### **Resonance production Comparison with other calculations**

- many calculations are devoted to exlusive pion production
- we compare our results with those of
  - S. Ahmad et al., nucl-th/0603001



the results seem to be in agreement

#### **Comparison with experimental data**

preliminary results

• from our results and standard isospin analysis we can estimate the cross section for 1  $\pi^+$ 

$$\begin{aligned} \mathcal{A}(\nu_{\ell} + p \to \ell^{-} + p + \pi^{+}) &= \mathcal{A}_{3} \\ \mathcal{A}(\nu_{\ell} + n \to \ell^{-} + n + \pi^{+}) &= \frac{1}{3}\mathcal{A}_{3} + \frac{2\sqrt{2}}{3}\mathcal{A}_{1} \\ \mathcal{A}(\nu_{\ell} + n \to \ell^{-} + p + \pi^{0}) &= -\frac{\sqrt{2}}{3}\mathcal{A}_{3} + \frac{2}{3}\mathcal{A}_{1} \end{aligned}$$

 $\downarrow$ 

 $A_3$  is the amplitude for the isospin 3/2 state of the  $\pi N$  system (predominantly the  $\Delta$ )  $A_1$  is the amplitude for the isospin 1/2 state.

$$\sigma_{\pi^+} = \frac{10}{9} \sigma_{\Delta^{++}} + \frac{8}{9} (b_1 \sigma_{P_{11}} + b_2 \sigma_{D_{13}} + b_3 \sigma_{S_{11}})$$

 $b_i$  are branching ratios for  $\pi^+$ 

#### **Comparison with experimental data**

preliminary results

- from our results and standard isospin analysis we can estimate the cross section for 1  $\pi^+$
- the MiniBooNE collaboration has presented a preliminary measurement of

$$R = \frac{\sigma_{\pi^+}}{\sigma_{CCQE}}$$

in  $(\nu_{\mu}\,,^{12}{\it C})$  production





• theoretical curves underestimate the (preliminary) data. Effects due to non-resonant background and coherent pion production could be sizeable

#### **Conclusions and outlooks**

- the  $\nu$ -nucleus cross-sections in IA regime with realistic spectral function have been discussed
- at  $E_{\nu} \sim \mathcal{O}(1 \text{ GeV})$  the quasi-elastic x-section with spectral function is lower than the widely applied Fermi Gas models; the inclusion of Pauli blocking further reduces the cross section
- the contribution of  $\triangle$ -production is not negligible
- other resonances have been added, resulting in a  ${\sim}10\%$  contribution to the resonance production cross section
- our results seem to be in agreement with other calculations
- future prospects:
  - inclusion of exclusive channels, both in the QE and resonance production regions
  - proper traitment of FSI (for differential CC x-sect and  $\pi$ -rescattering)
  - inclusion of non-resonant pion production

# backup slides

### **Elastic Structure Functions and Form Factors**

• structure functions

$$W_{1} = 2 \left[ -\frac{q^{2}}{2} (F_{1} + F_{2})^{2} + \left( 2 m_{N}^{2} - \frac{q^{2}}{2} \right) F_{A}^{2} \right]$$

$$W_{2} = 4 \left[ F_{1}^{2} - \left( \frac{q^{2}}{4 m_{N}^{2}} \right) F_{2}^{2} + F_{A}^{2} \right]$$

$$W_{3} = -4 (F_{1} + F_{2}) F_{A}$$

$$W_{4} = -2 \left[ F_{1} F_{2} + \left( 2 m_{N}^{2} + \frac{q^{2}}{2} \right) \frac{F_{2}^{2}}{4 m_{N}^{2}} + \frac{q^{2}}{2} F_{P}^{2} - 2 m_{N} F_{P} F_{A} \right]$$

$$W_{5} = \frac{W_{2}}{2}$$

• form factors

$$\tau = \frac{q^2}{4 m_N^2} \quad G_E = \frac{1}{\left(1 - \frac{q^2}{M_V^2}\right)^2} \quad G_M = 4.71 \, G_E$$

$$F_1 = \frac{1}{1 - \tau} \left(G_E - \tau \, G_M\right) \quad F_2 = \frac{1}{1 - \tau} \left(-G_E + G_M\right)$$

$$F_A = -\frac{1.26}{\left(1 - \frac{q^2}{M_A^2}\right)^2} \quad F_p = -\frac{1.28}{\left(1 - \frac{q^2}{0.14}\right)^2} \frac{F_A}{-1.27}$$

#### **Anelastic Structure Functions and Form Factors**

- structure functions are complicated functions of form factors
- detailed formulae in: Lalakulich et al., Phys.Rev. D71, 074003 (2005) and Lalakulich et al., Phys.Rev. D74, 014009 (2006)
- form factors for  $\Delta$ -production

$$\begin{aligned} f_{\pi} &= 0.97 \, m_{\pi} \quad g_{D} = 15.3 \quad M_{V} = 0.84 \quad M_{A} = 1.05 \\ C_{3}^{V}(0) &= 1.95 \quad C_{3}^{V}(q^{2}) = \frac{C_{3}^{V}(0)}{\left(1 - \frac{q^{2}}{M_{V}^{2}}\right)^{2}} \frac{1}{\left(1 - \frac{q^{2}}{4M_{V}^{2}}\right)} \\ C_{4}^{V}(q^{2}) &= -C_{3}^{V}(q^{2}) \frac{m_{N}}{W} \quad C_{5}^{V} = 0 \\ C_{5}^{A}(0) &= \frac{f_{\pi} \, g_{D}}{\sqrt{3}} \quad C_{3}^{A} = 0 \quad C_{5}^{A}(q^{2}) = \frac{C_{5}^{A}(0)}{\left(1 - \frac{q^{2}}{M_{A}^{2}}\right)^{2}} \frac{1}{\left(1 - \frac{q^{2}}{3M_{A}^{2}}\right)} \\ C_{4}^{A}(q^{2}) &= -\frac{C_{5}^{A}}{4} \quad C_{6}^{A}(q^{2}) = C_{5}^{A} \frac{m_{N}^{2}}{-q^{2} + m_{\pi}^{2}} \end{aligned}$$

#### Elastic interactions Comparison with other calculations

different approach followed by C. Maieron et al., nucl-th/0303075
 Impulse approximation to describe QE scattering +
 Relativistic Shell Model to describe the bound nucleon +
 Relativistic Optical Potential (ROP) or Relativistic Mean Field (RMF) to account for FSI



good agreement with cross sections including SF + Pauli blocking

# $(\nu_{\mu},^{12}\,C) \label{eq:multiplicative}$ cross section of muon neutrinos on Carbon



#### Elastic interactions Comparison with other calculations

• the extension at higher energies: S. Ahmad et al., nucl-th/0603001



- we tried with two different vector and axial vector masses
- some discrepancy is found at energies  $E_{\nu} > 0.5$  GeV not completely understood

#### Some detail of the spectral function calculation

$$\langle 0|J^{\mu}|X\rangle = \frac{m}{\sqrt{\mathbf{p}_{\mathcal{R}}^2 + m^2}} \langle 0|\mathcal{R}, \mathbf{p}_{\mathcal{R}}; \mathbf{N}, -\mathbf{p}_{\mathcal{R}}\rangle \\ \times \sum_{i} \langle -\mathbf{p}_{\mathcal{R}}, N|j_i^{\mu}|x, \mathbf{p}_x\rangle ,$$

$$W_{A}^{\mu\nu} = \sum_{x,\mathcal{R}} \int d^{3}p_{\mathcal{R}} d^{3}p_{x} |\langle 0|\mathcal{R}, \mathbf{p}_{\mathcal{R}}; \mathbf{N}, -\mathbf{p}_{\mathcal{R}} \rangle|^{2} \frac{m}{E_{\mathbf{p}_{\mathcal{R}}}}$$
$$\times \sum_{i} \langle -\mathbf{p}_{\mathcal{R}}, \mathbf{N} | j_{i}^{\mu} | x, \mathbf{p}_{x} \rangle \langle \mathbf{p}_{x}, x | j_{i}^{\nu} | \mathbf{N}, -\mathbf{p}_{\mathcal{R}} \rangle$$
$$\times \delta^{(3)} (\mathbf{q} - \mathbf{p}_{\mathcal{R}} - \mathbf{p}_{x}) \delta(\nu + E_{0} - E_{\mathcal{R}} - E_{x}),$$

$$P(\mathbf{p}, E) = \sum_{\mathcal{R}} |\langle 0|\mathcal{R}, -\mathbf{p}; \mathbf{N}, \mathbf{p} \rangle|^{2} \\ \times \delta(E - m + E_{0} - E_{\mathcal{R}}) ,$$

in which  $E_0$  is the energy of the initial hadronic state. The other variables have been already

#### introduced

• to evaluate the relative importance of QE and  $\Delta$ -production at  $\beta$ -Beams, we compute (fluxes from J. Burguet-Castell et at., Nucl.Phys.B695:217-240,2004)

$$\bar{\sigma}_{i}(\gamma) = \frac{\int dE_{\nu} \Phi_{\nu_{e}}(E_{\nu}, \gamma) \sigma_{i}(E_{\nu})}{\int dE_{\nu} \Phi_{\nu_{e}}(E_{\nu}, \gamma)} \qquad i = QE, \Delta$$

and plot the ratio  $\sigma_\Delta/\sigma_{QE}$  as a function of the boost factor  $\gamma$ 

